



Algorithms for the mitigation of space weather threats at low latitudes, contributing to the extension of EGNOS over Africa

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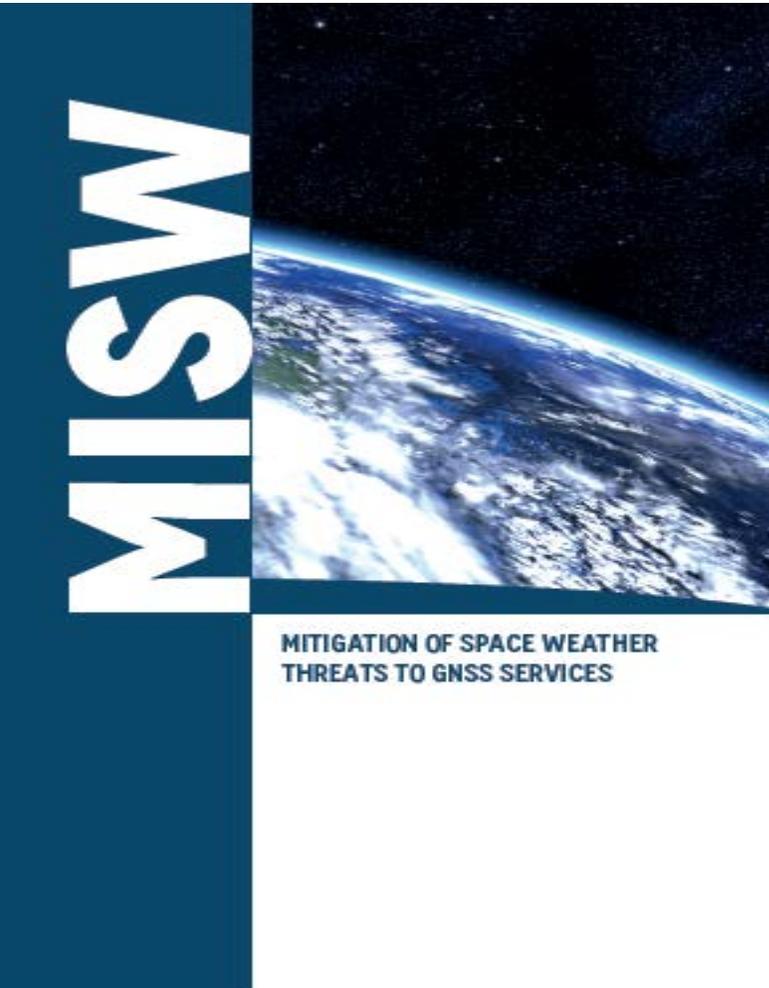
(1) University of Bath, UK, (2) Thales Alenia Space Italy, Roma and Milano, ITALY, (3) University of Nottingham, UK, (4) Politecnico di Torino, ITALY, (5) Istituto Nazionale di Geofisica e Vulcanologia, ITALY, (6) EISCAT Scientific Association, SWEDEN, (7) European Joint Research Centre, Institute for Security and Protection of Citizens, BELGIUM, (8) Technical University of Denmark - Space, DENMARK, (9) Space Research Centre, POLAND, (10) University of Zagreb, CROATIA, (11) Met Office, Exeter, UK.

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<http://misw.info/>



The challenge



WHY MISW?

- Space Weather affects many modern technologies that we take for granted
- MISW tackles research challenges associated with GNSS and Space Weather
- Bring practical solutions to European Industry

USERS

- GNSS Users rely on positioning accuracy but this may not be adequate for all applications
- Safety critical applications also need integrity of the GNSS positioning

SBAS

- SBAS systems (EGNOS in Europe) gather information that allows some mitigation of Space Weather events
- However, SBAS systems are not yet able to work in the most challenging regions and as consequence Space Weather disturbances to the ionized upper atmosphere cause navigation errors that remain uncompensated

Ionospheric Threats

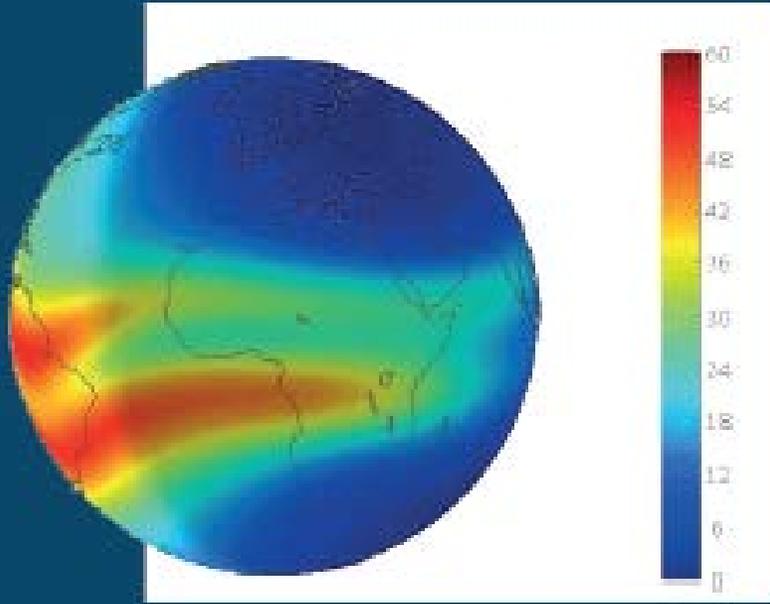


Illustration of global ionosphere at solar max. Colour contours show vertical TEC where 60 TECU are equivalent to delays on the GPS L1 signal of about 10 m vertical.

IONIZATION GRADIENTS

- Ionization gradients (in particular over low latitude regions) are much more structured than over Europe with high TEC gradients

SCINTILLATION

- Propagation of radio signals through ionization gradients lead to scintillation (Fluctuation of signal phase and amplitude)
- Scintillation is a serious threat to GNSS as it can disrupt receivers operation and service entirely by means of C/N_0 fading and loss of lock

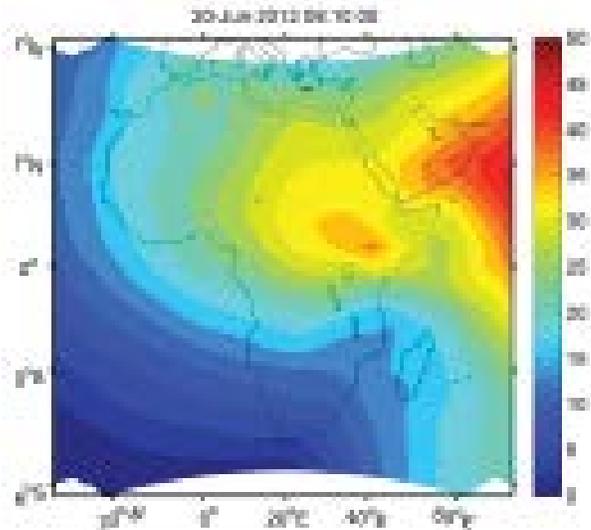
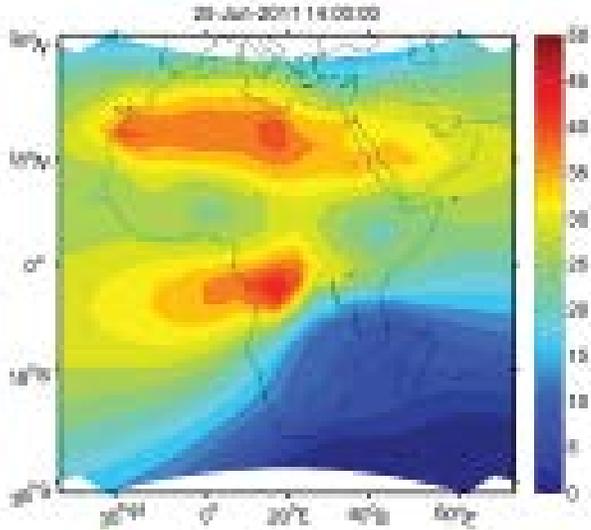
MISW Solutions

EGNOS EXTENSION

- EGNOS extends from Scandinavia in the North to Africa to the South
- These regions experience strong gradients in delay and break up of signals from scintillation
- These issues represent a technology barrier to the expansion of EGNOS geographically

MISW OBJECTIVES

- Monitor and characterize iono effects at low, mid, high latitudes
- Develop algorithms against Space Weather vulnerabilities at Receiver level and at System level
- Enable extension of EGNOS to Africa
- Devise recommendations on best practices for GNSS services with reference to Space Weather



MISW Partners

	University of Bath, UK
	Thales Alenia Space ITALY, FRANCE
	University of Nottingham, UK
	Polytechnic of Torino, ITALY
	National Institute of Geophysics and Vulcanology, ITALY (INGV)
	EISCAT Scientific Association, SWEDEN
	European Joint Research Centre Institute for Security and Protection of Citizens - BELGIUM
	Danish Technological University - Space
	Space Research Centre - POLAND
	University of Zagreb - CROATIA
	MET Office, UK

PARTNERS

- Under the lead of University of Bath the MISW partners include major european institutions and industry involved in GNSS and Space Weather study

STAKEHOLDERS

- Relevant entities regulating SBAS services
- Experts in Ionospheric studies
- GNSS User Communities
- GNSS Industry
- GNSS Service Providers

SUPPORTING PARTNERS

- ASECNA, SANSA, CIVIL AVIATION UNIVERSITY OF CHINA, NOAA, ESA, FAA

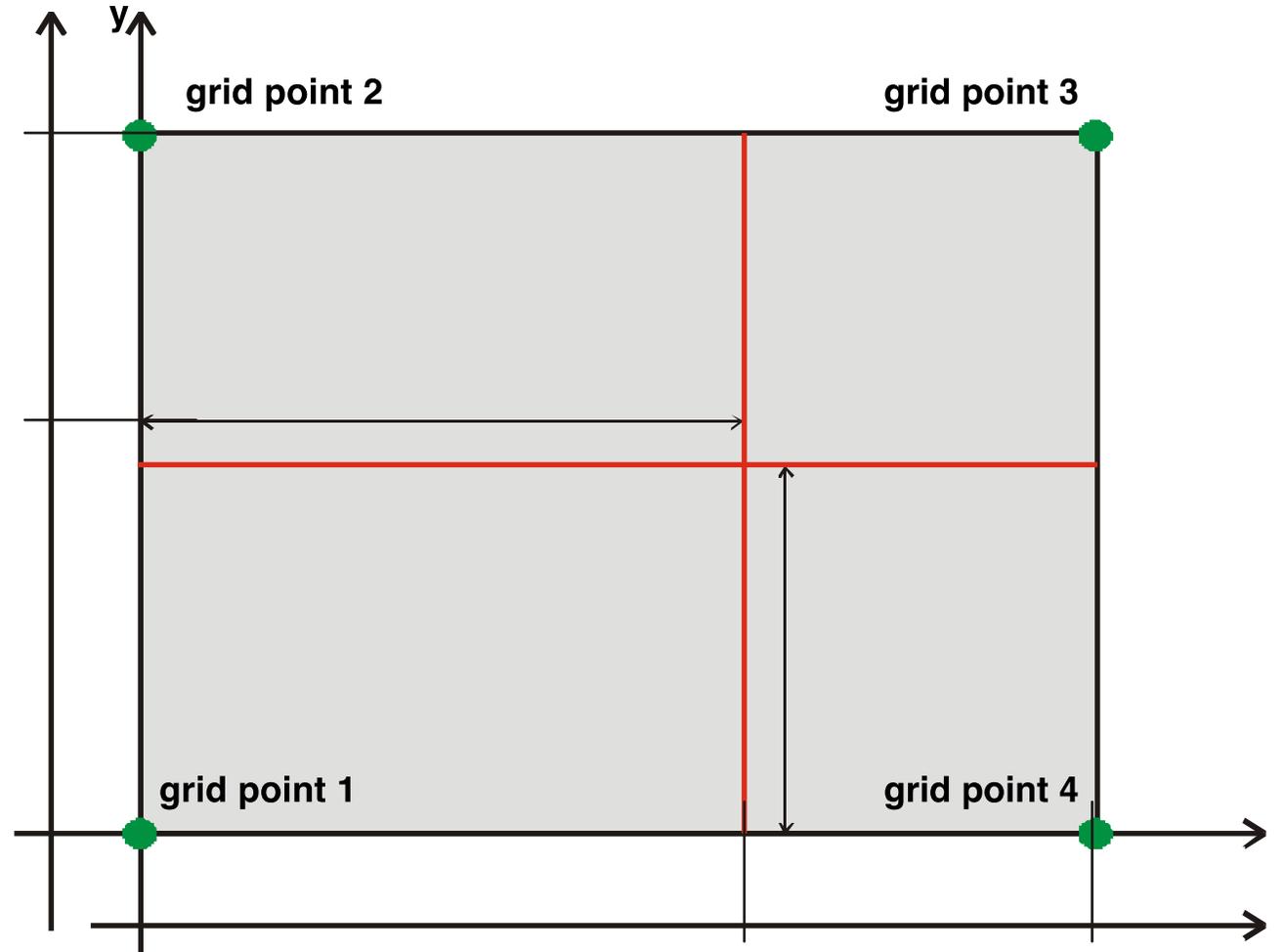
The concept of grid point corrections

How to calculate corrections to ionospheric delays

$$\tau_{vpp}(\Phi_{pp}, \lambda_{pp}) = \sum_{i=1}^4 W_i(x_{pp}, y_{pp}) \tau_{vi}$$

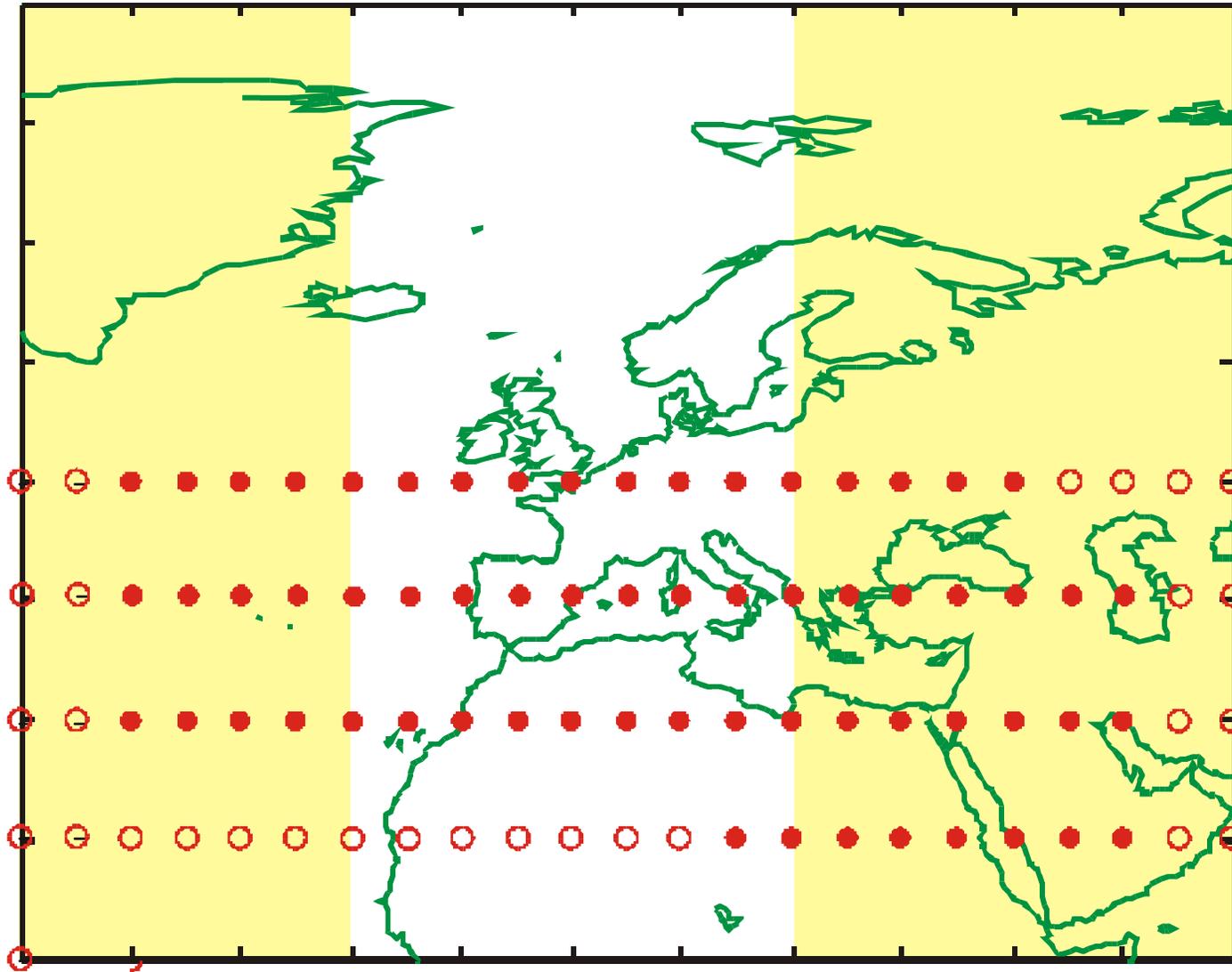
Credit: RTCA

with τ_{vpp} vertical ionospheric delay at pierce point
 τ_{vi} vertical ionospheric delay at grid points
 W_i weighting function



Example of ionospheric grid points

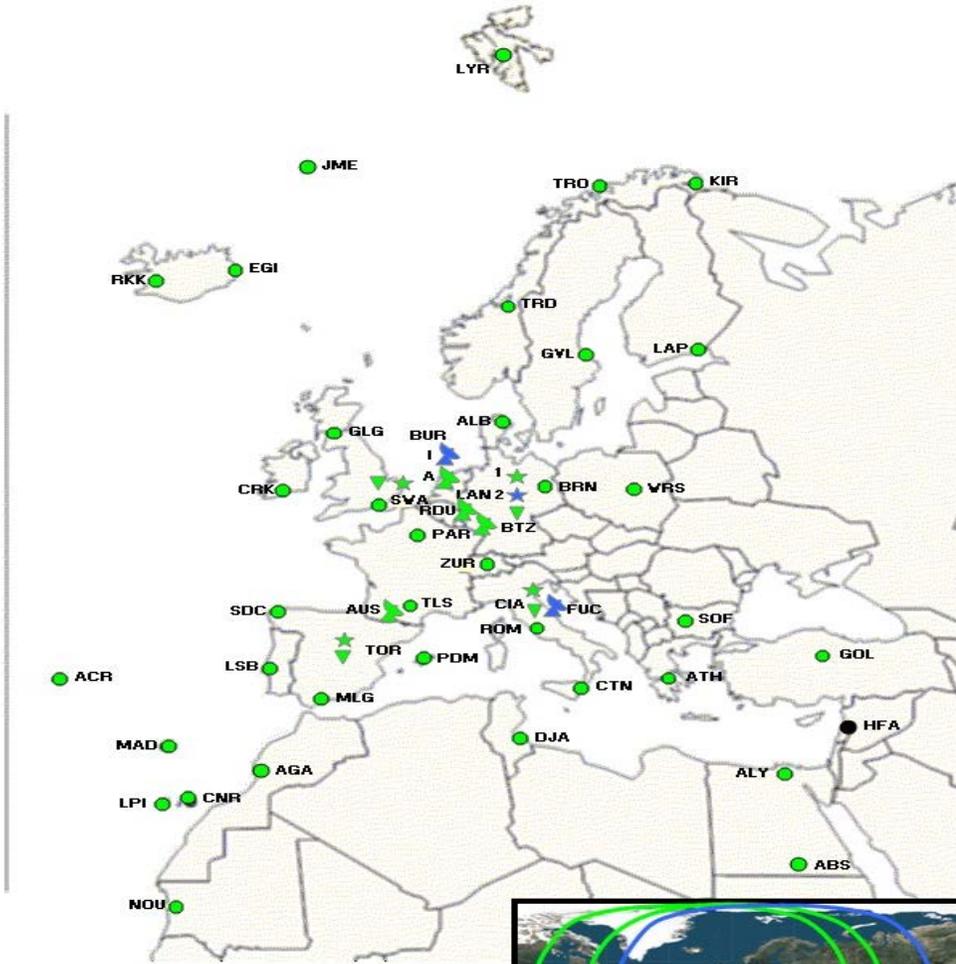
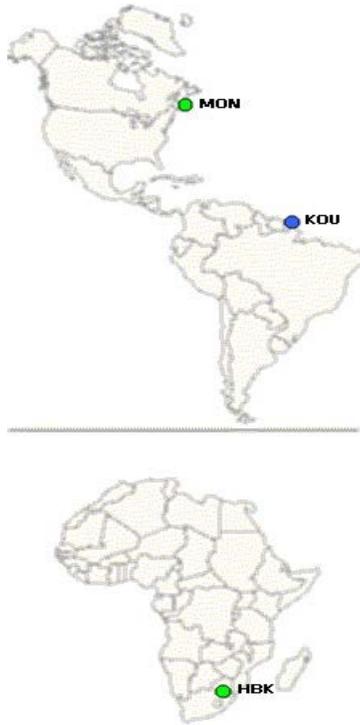
Credit: RTCA



EGNOS monitoring stations - courtesy ESSP



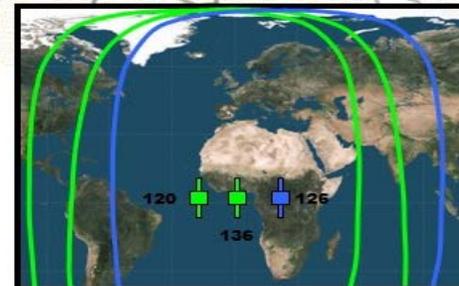
EGNOS System
- Segregation -
Since 2015/08/20 14:09



02/09/2015 © ESSP

Legend :

Assets:	Status:
○ RIMS	■ EGNOS-OP
☆ CPF	■ NEW EGNOS-TEST
▽ CCF	■ EGNOS-TEST ONLY
▲ NLES	● Not deployed (Location and name TBC)
□ GEO PRN	

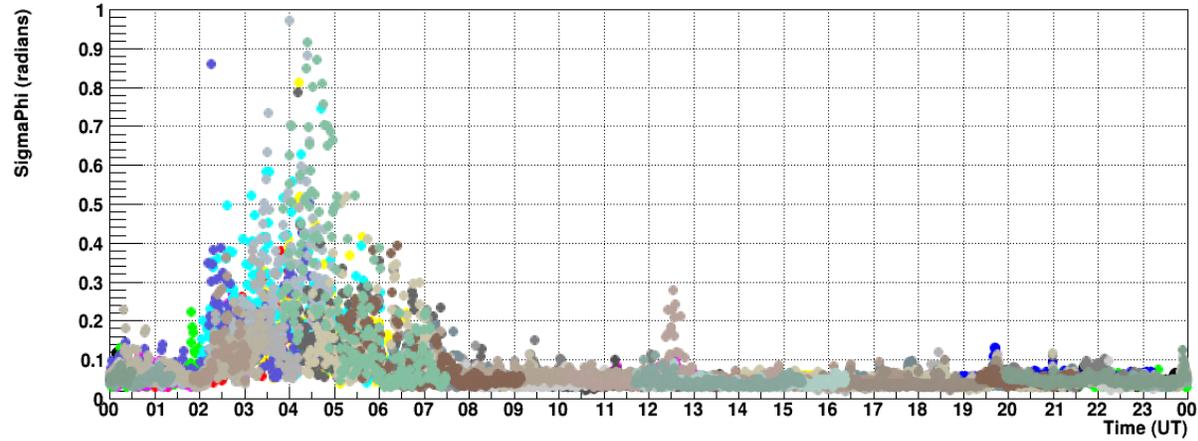


Credit: ESSP

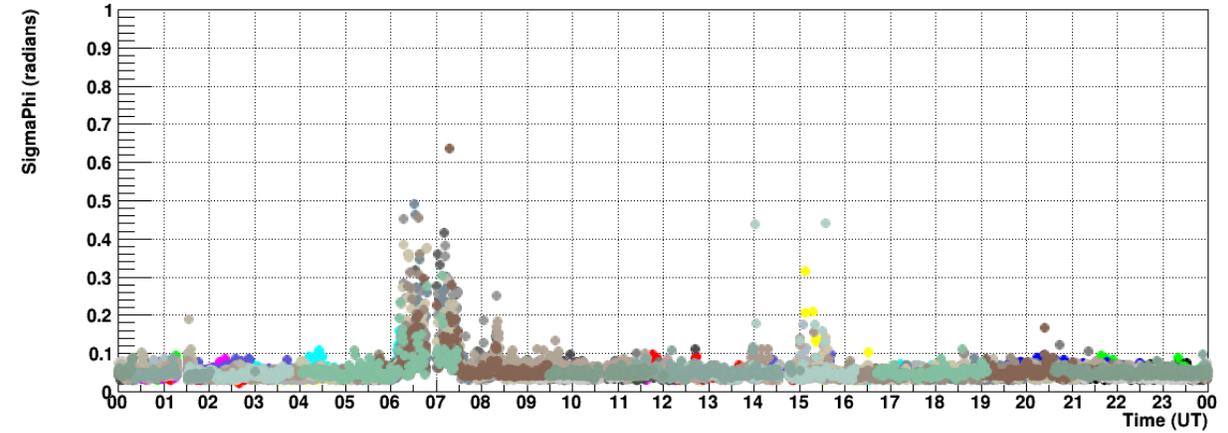
MISW Solutions: Ionospheric Scenarios

Rate of Change of TEC and Scintillation

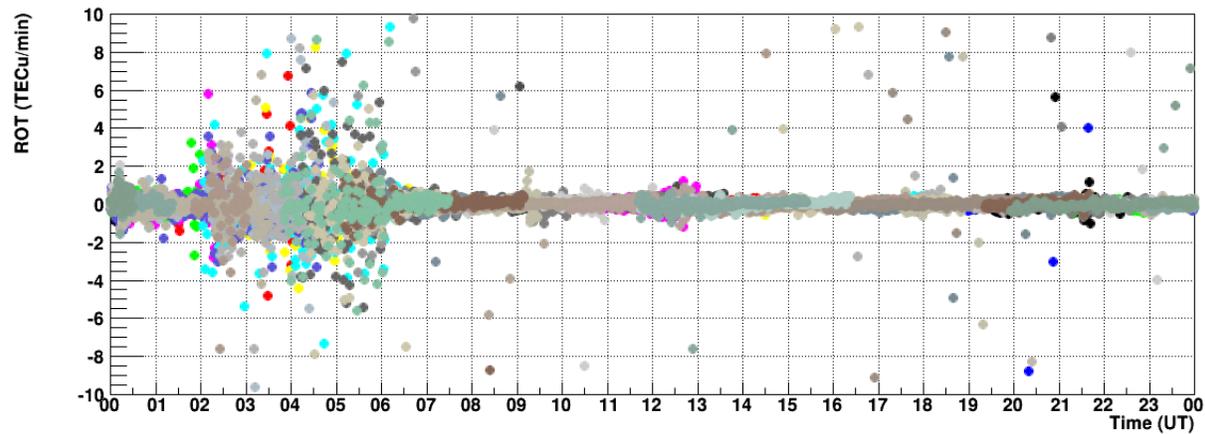
SigmaPhi index



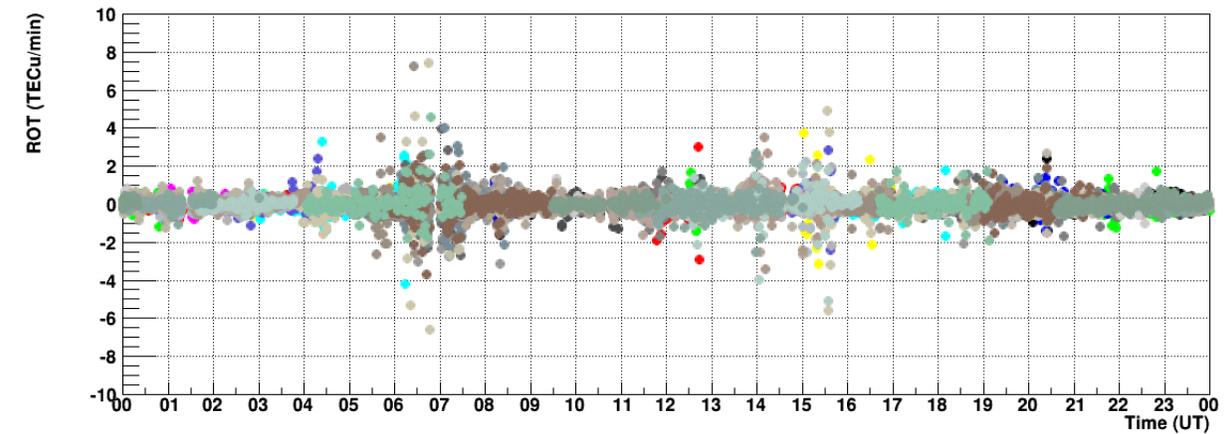
SigmaPhi index



ROT



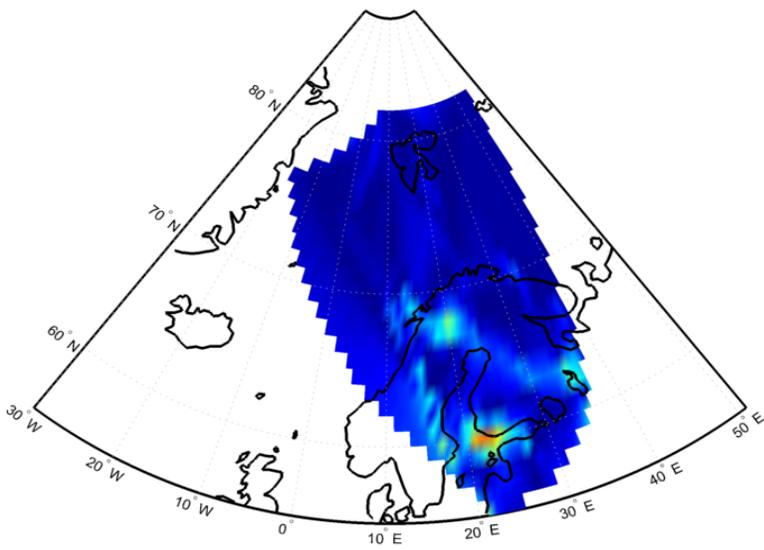
ROT



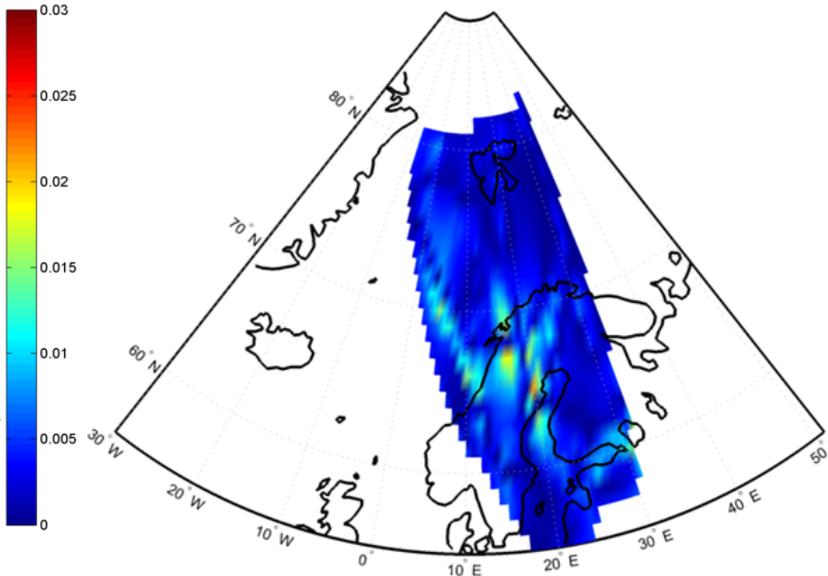
23 June 2015
Trondheim (63.42 N, 10.41 E)

23 June 2015
Ny Alesund (78.93 N, 11.06 E)

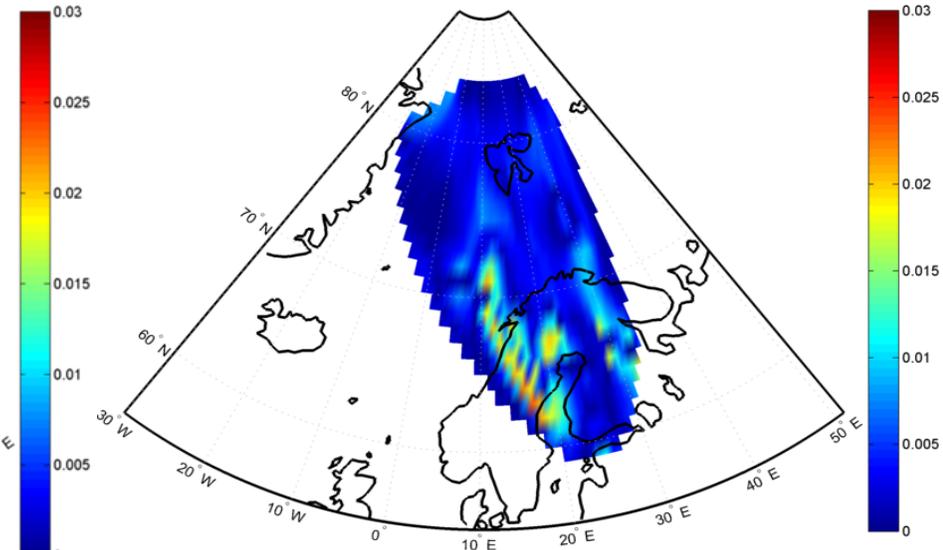
23-Jun-2015 02:00:00 TECU/km N-S



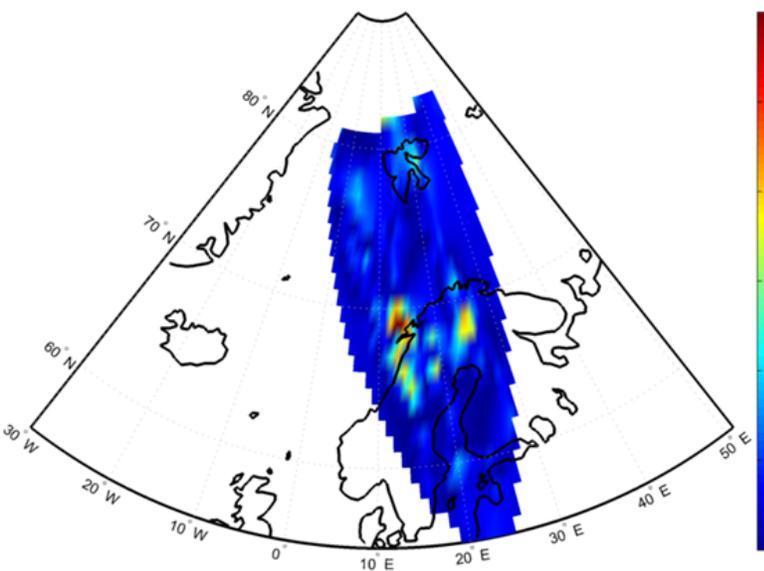
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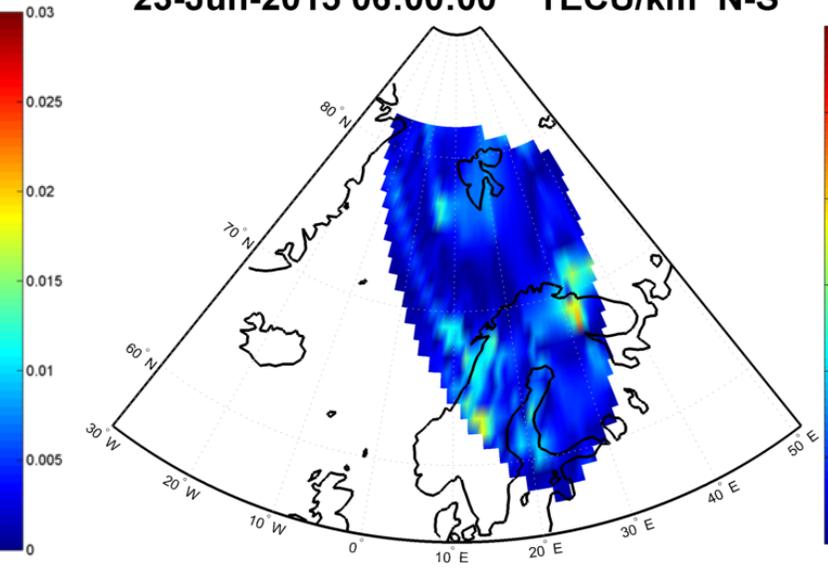
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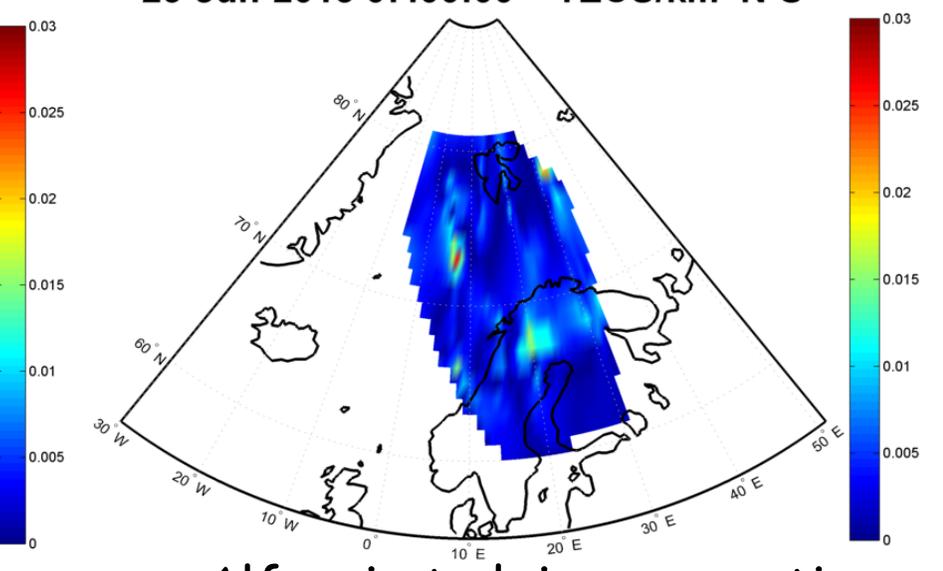
23-Jun-2015 05:00:00 TECU/km N-S



23-Jun-2015 06:00:00 TECU/km N-S



23-Jun-2015 07:00:00 TECU/km N-S

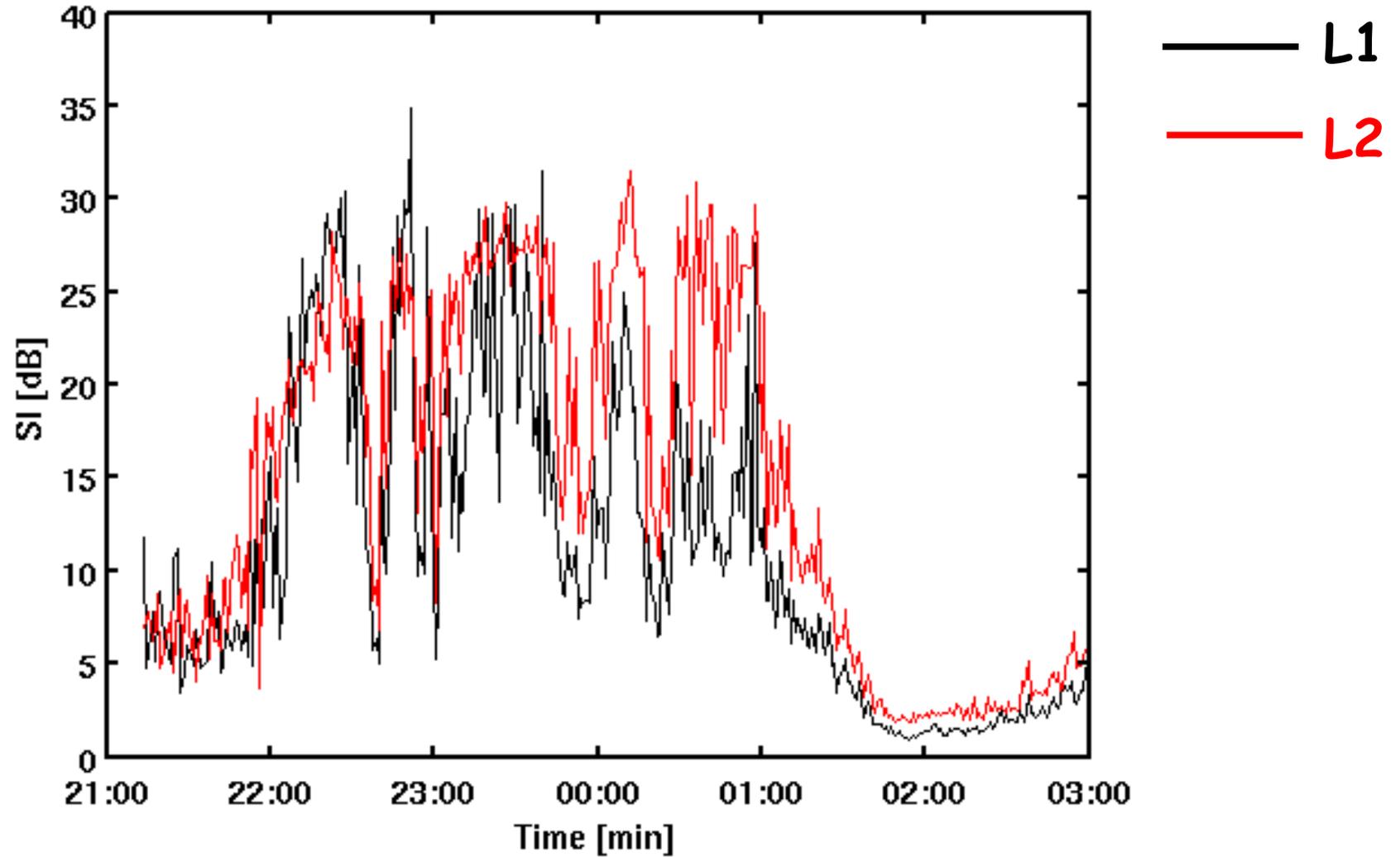


Alfonsi et al, in preparation

**An additional problem at low latitudes:
scintillation**

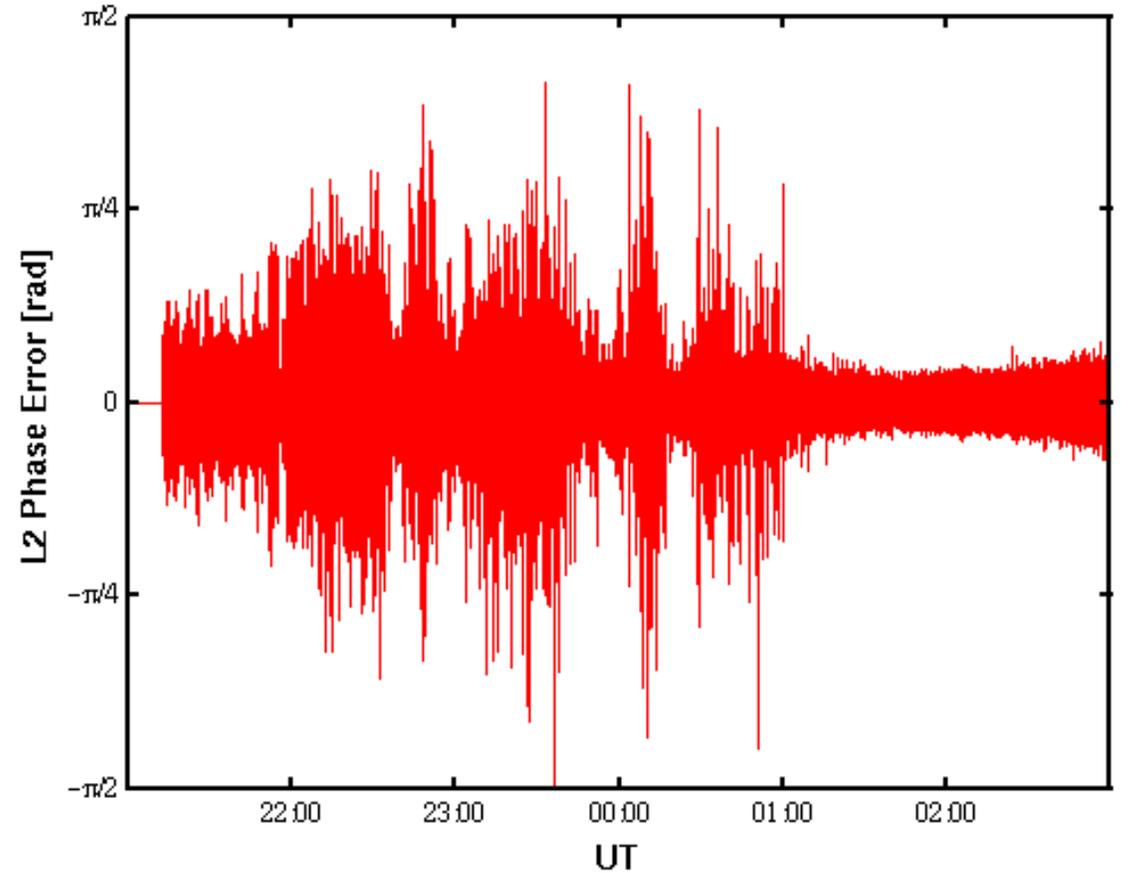
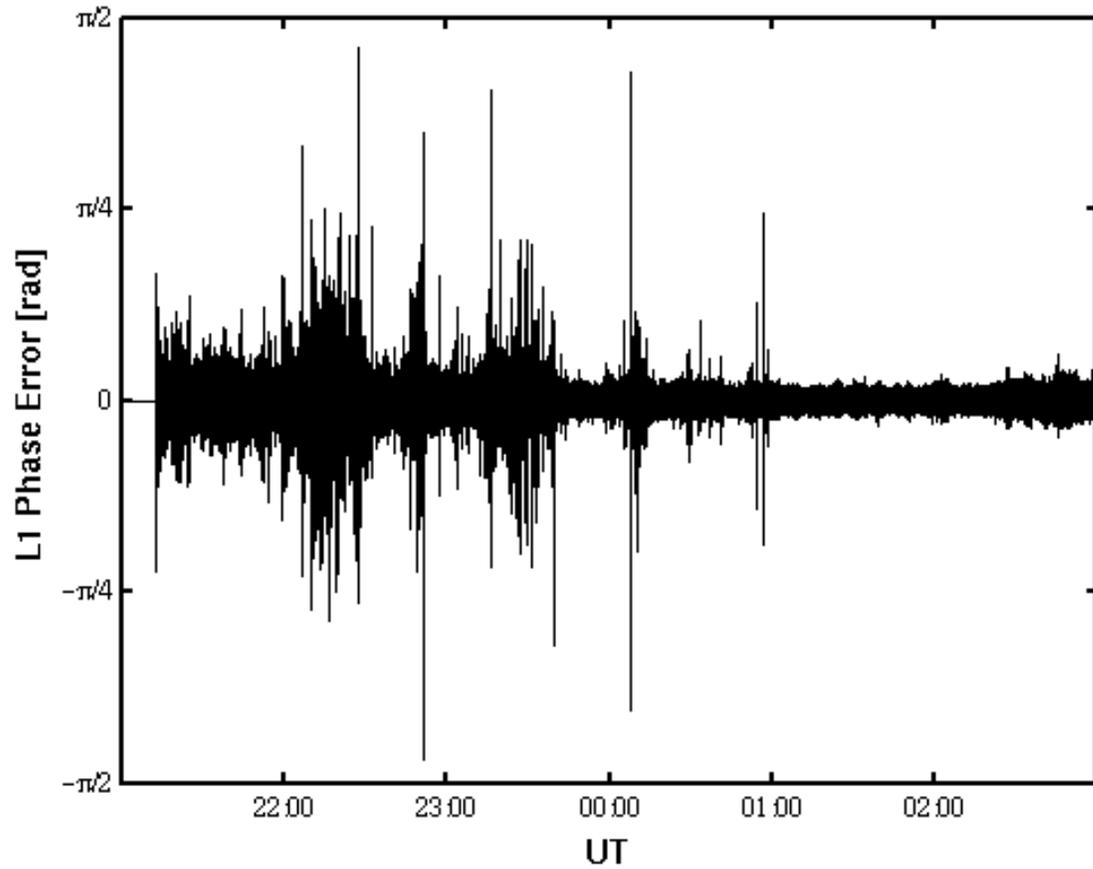
An additional problem: scintillation

10 March 2012
PRN31



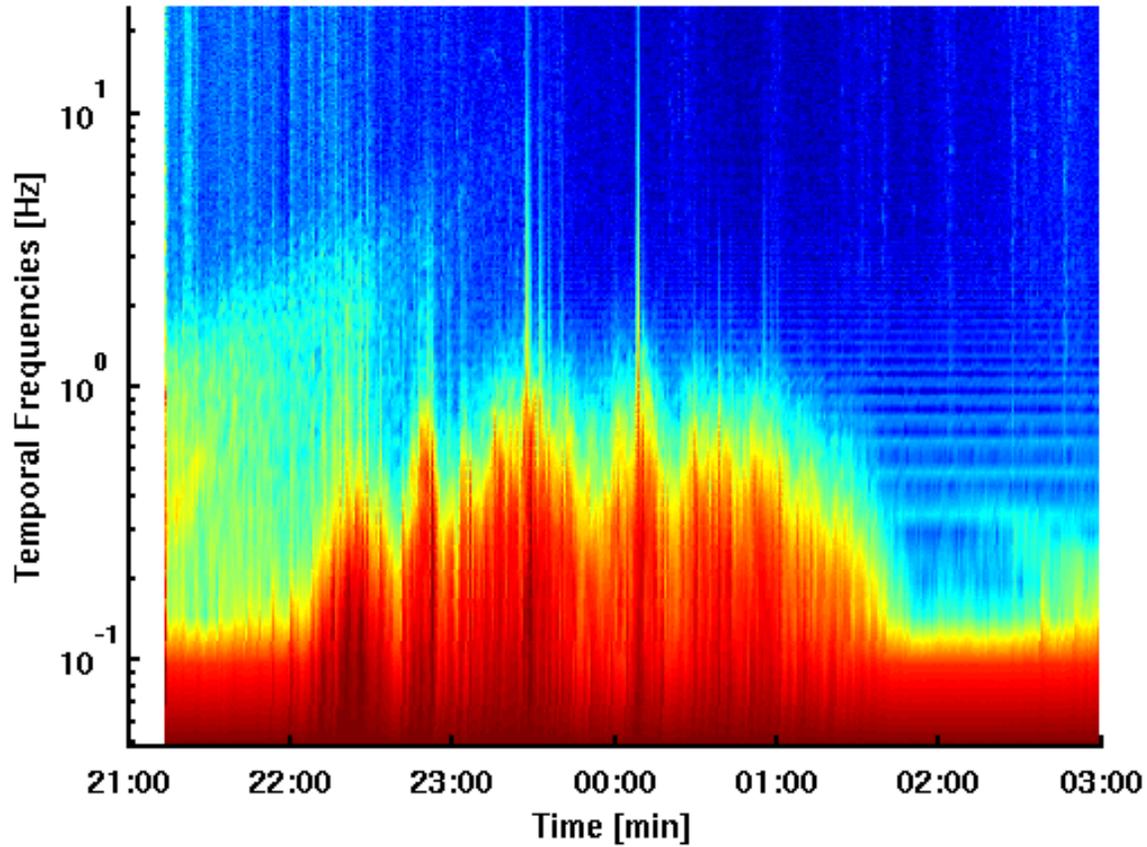
Forte et al, in preparation

An additional problem: scintillation

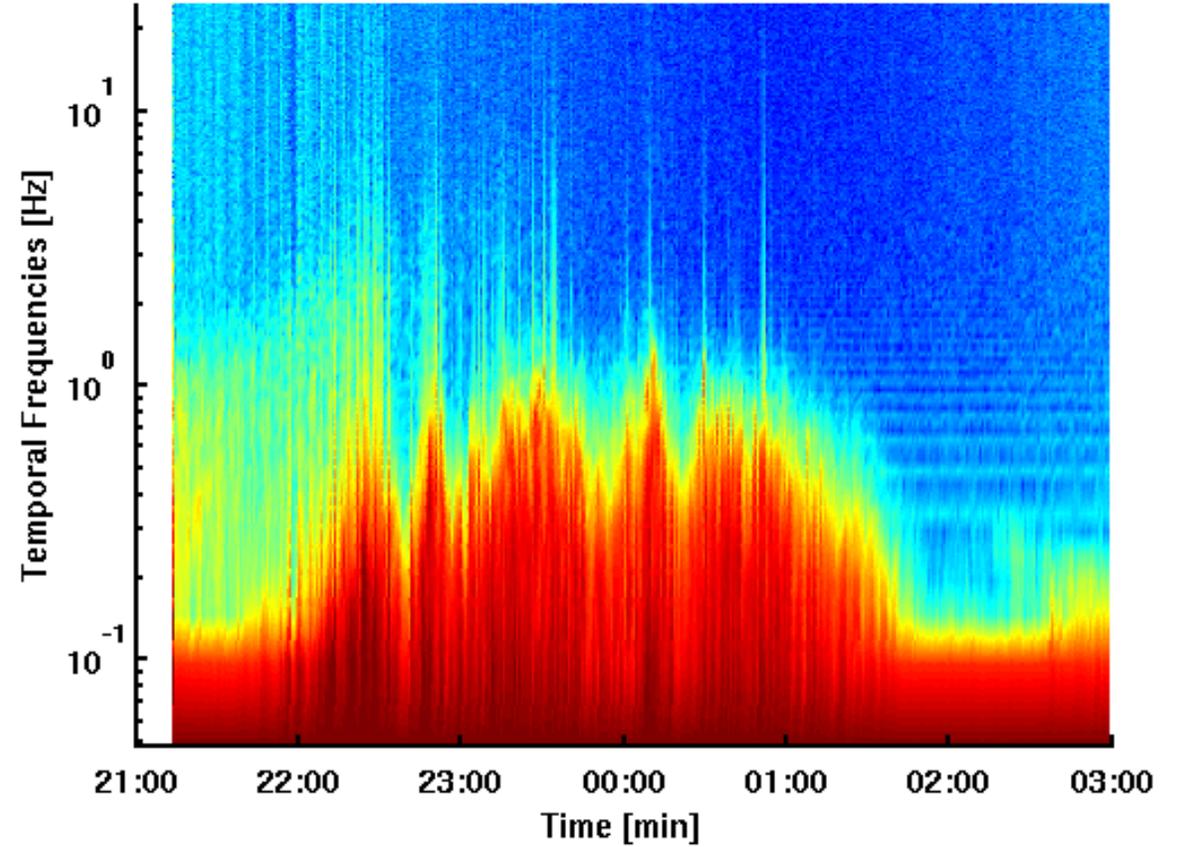


Forte et al, in preparation

An additional problem: scintillation



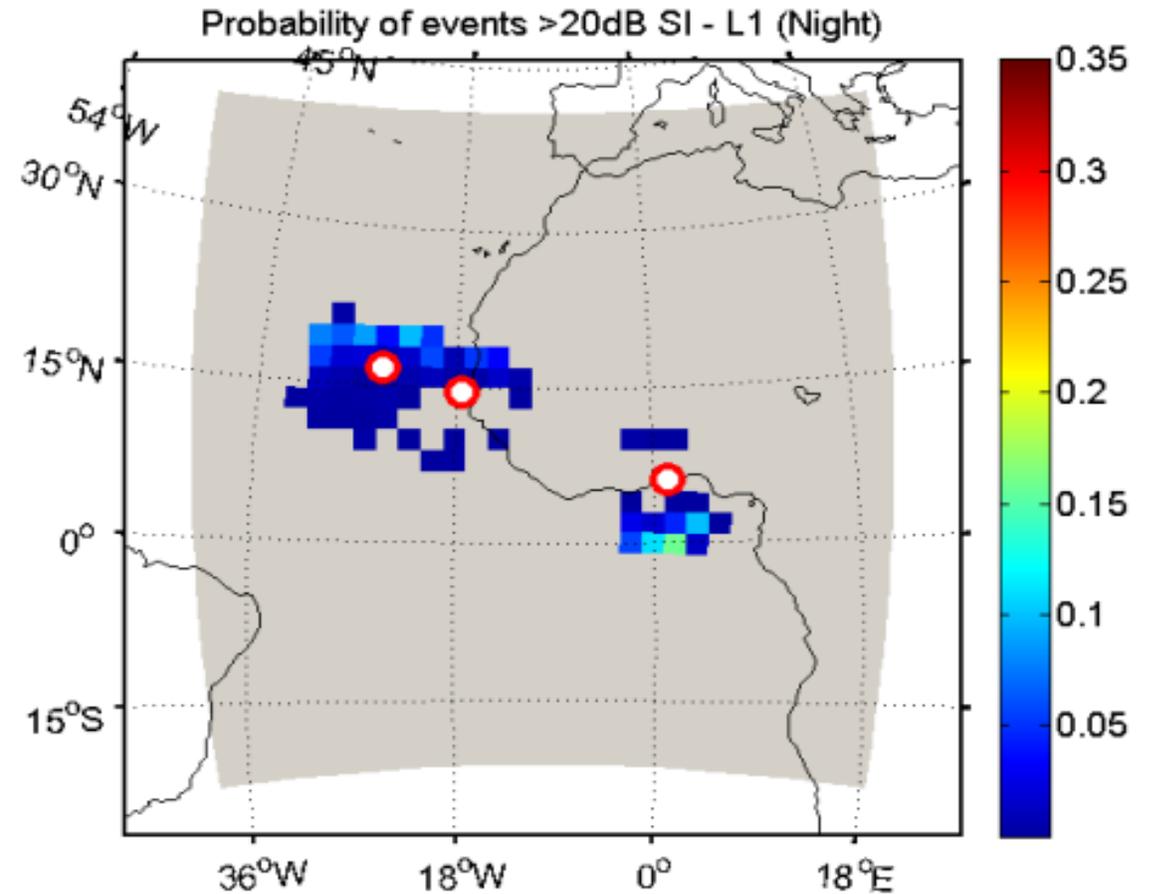
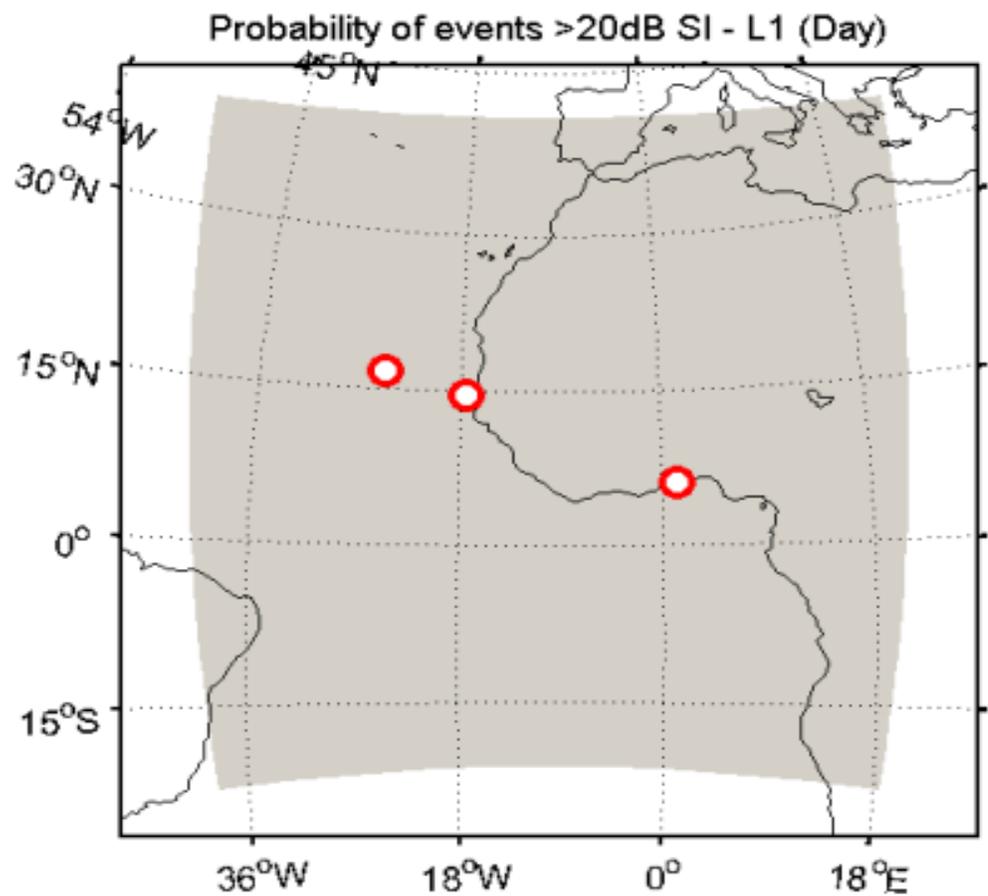
L1



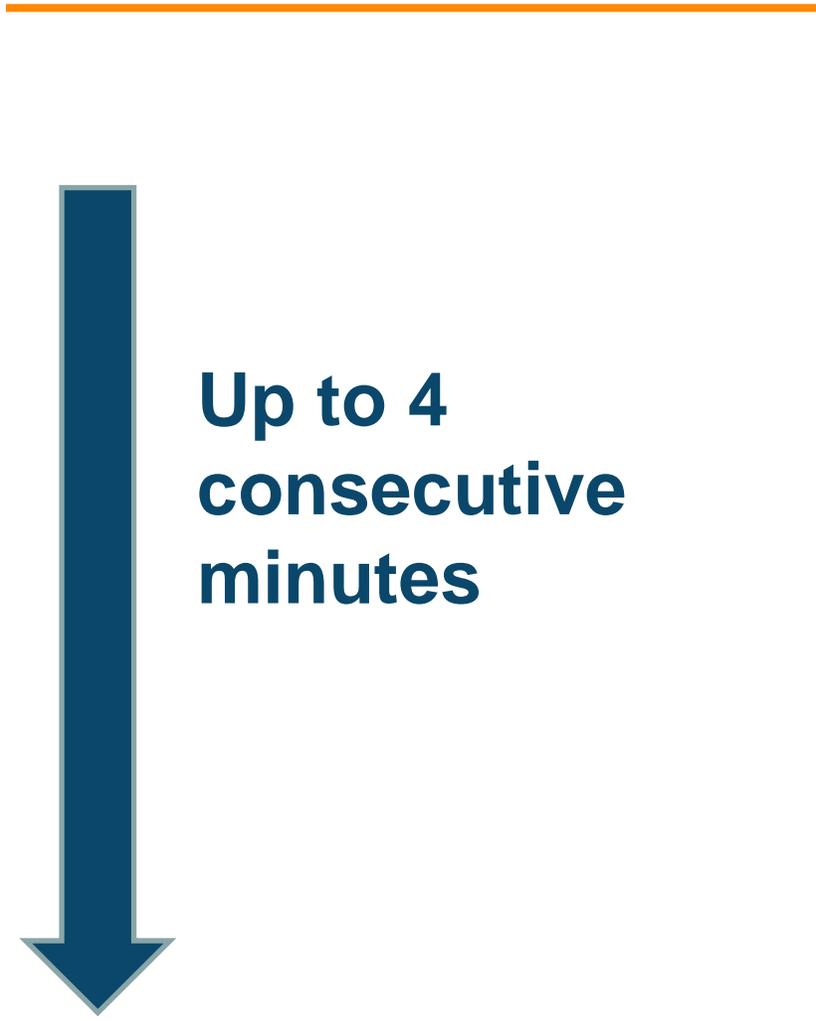
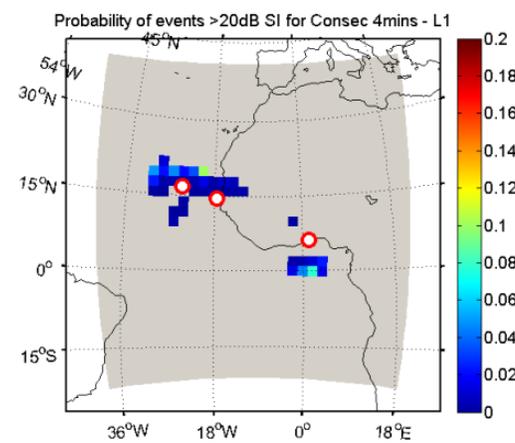
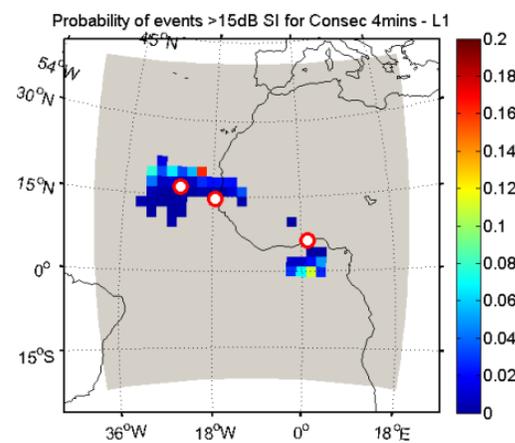
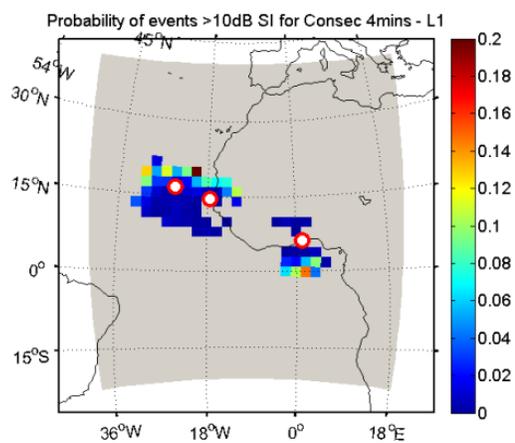
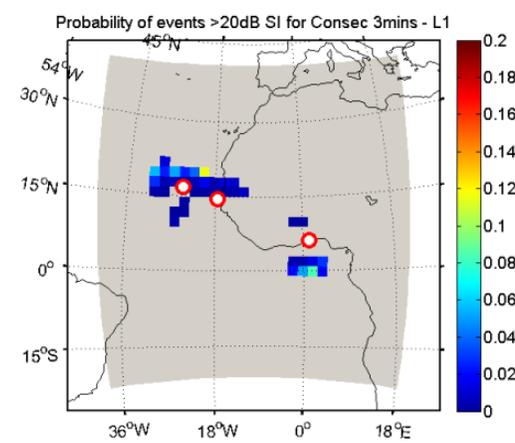
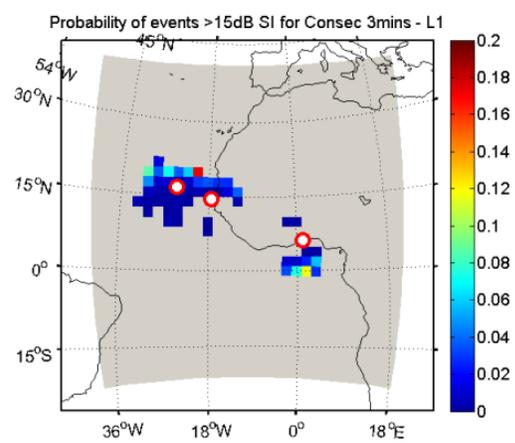
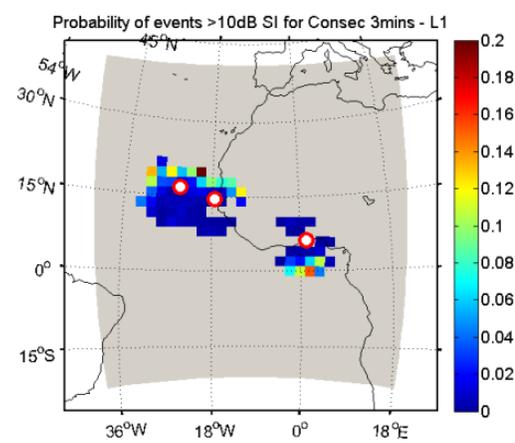
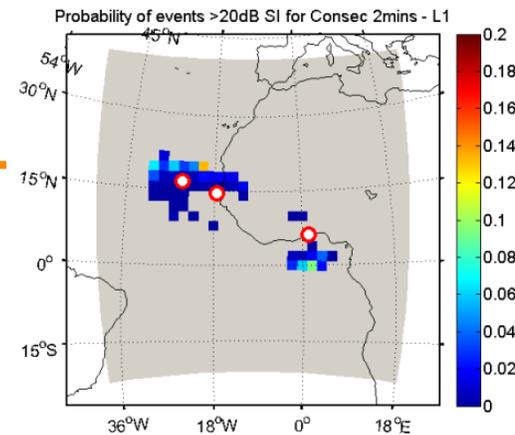
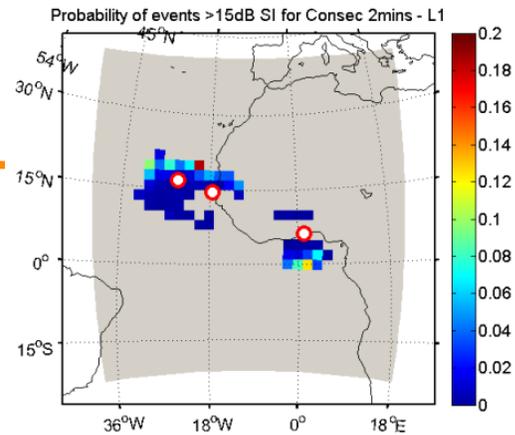
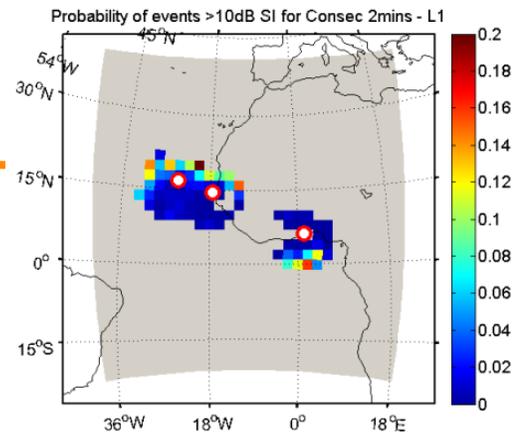
L2

Forte et al, in preparation

Scintillation: a night-time phenomenon



Forte et al, in preparation



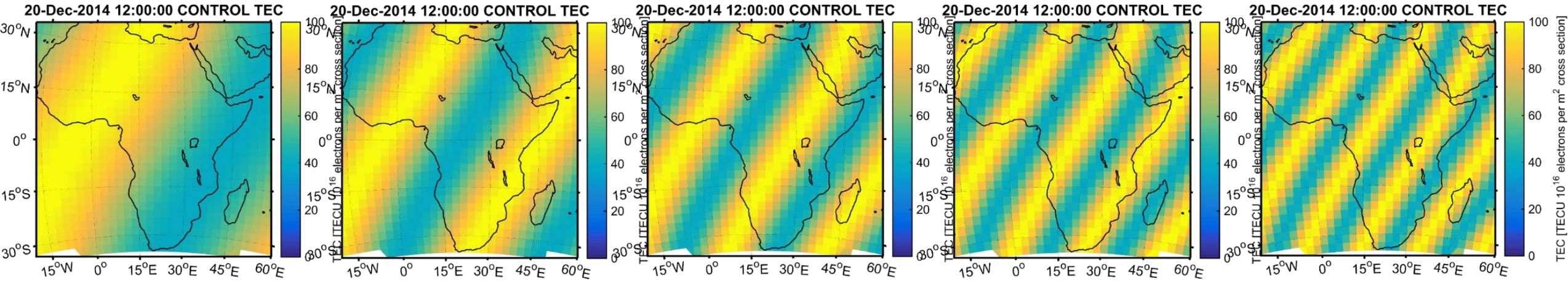
**Up to 4
consecutive
minutes**

Forte et al, in preparation

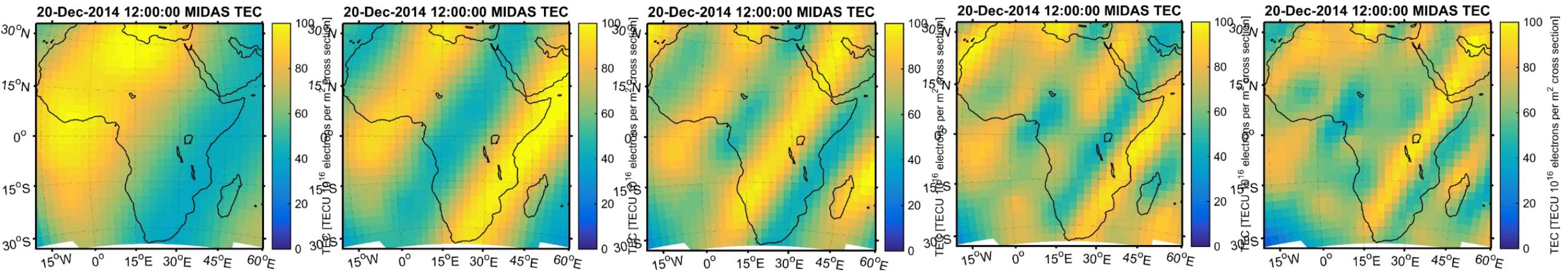
MISW Solutions: Understanding complexity and data requirement

1. Complexity

Simulated



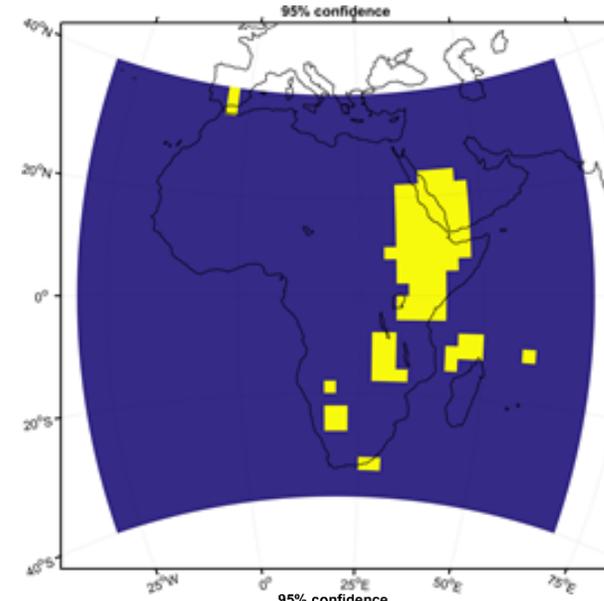
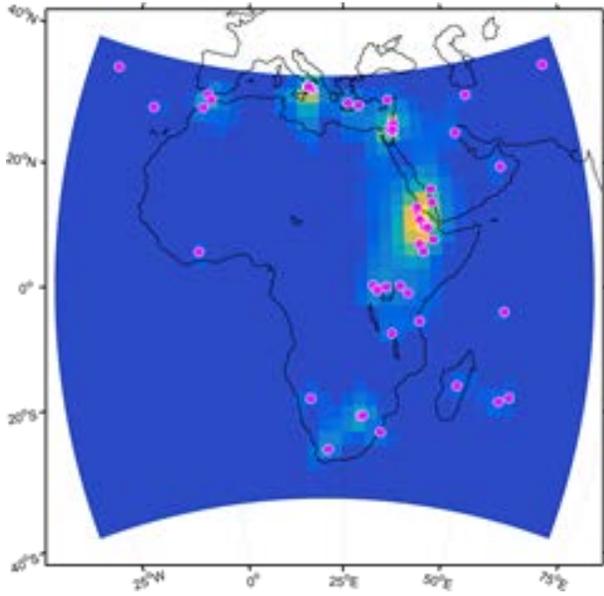
Reconstructed



Da Dalt et al, in preparation

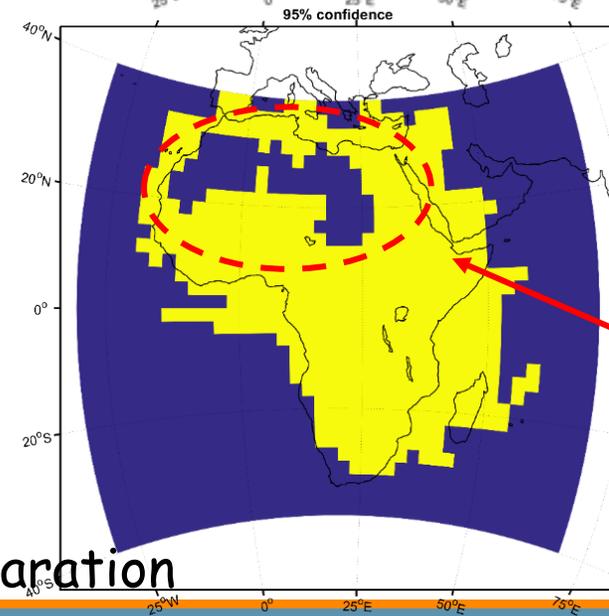
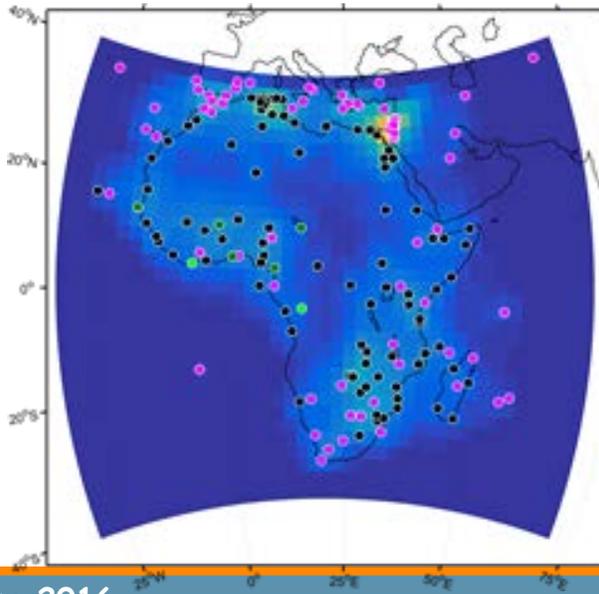
2. Data coverage

2009
Low data coverage



95% confidence

2015 +
simulated
receivers
High data coverage

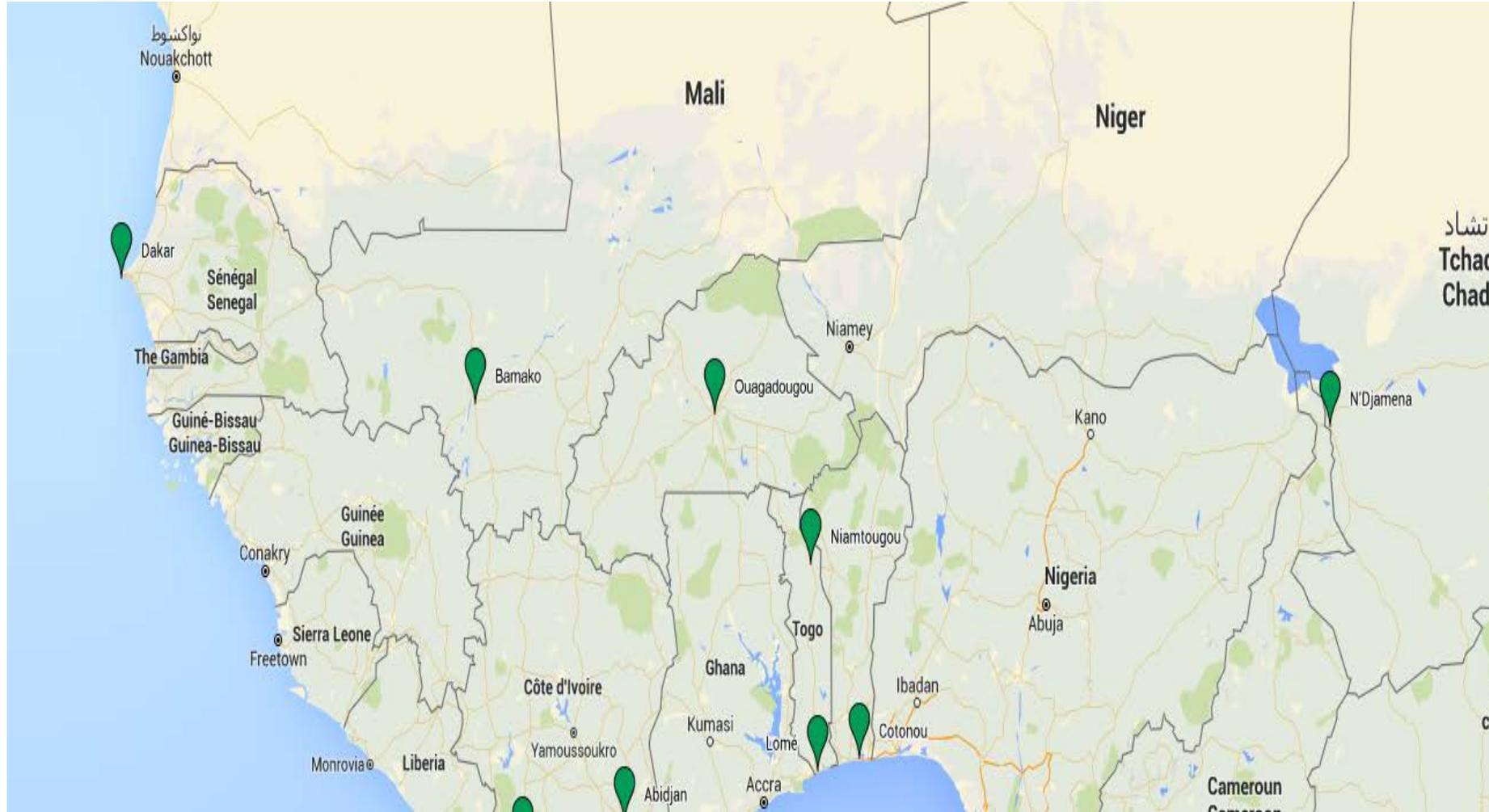


Sahara desert

Da Dalt et al, in preparation

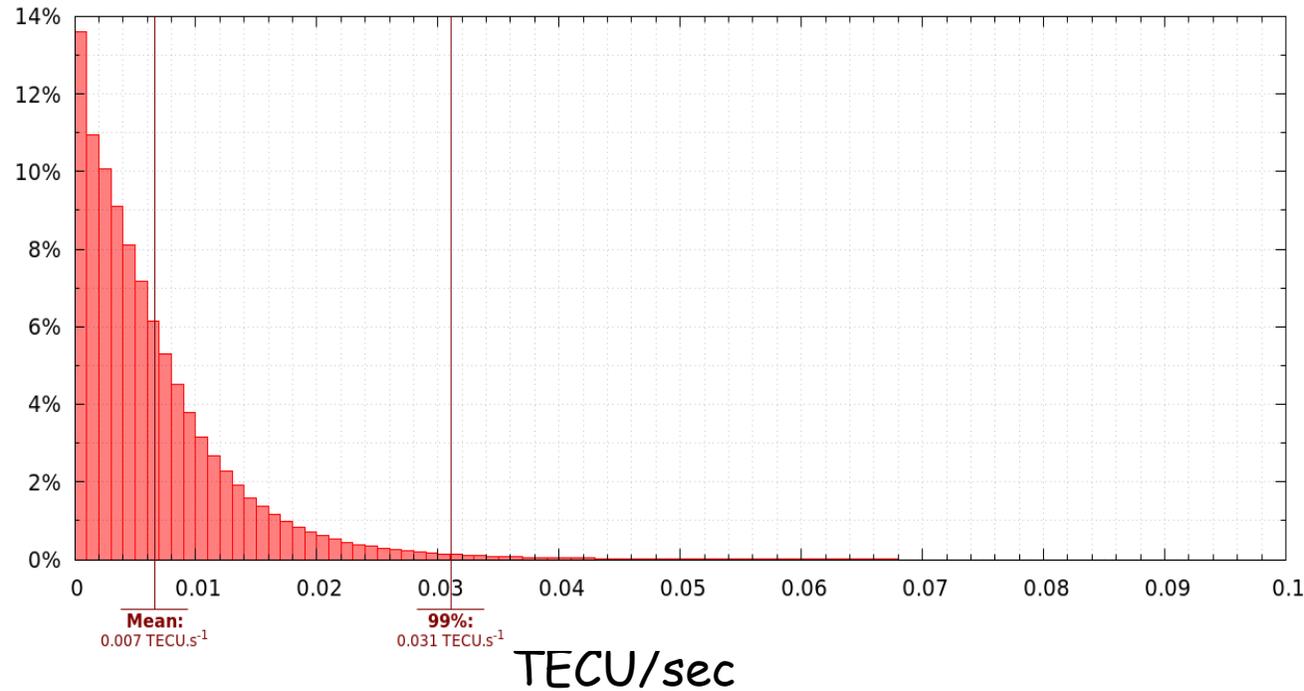
MISW Solutions: Modelling EGNOS performance over Africa under given scenarios

Data from SAGAIE network - courtesy of French CNES

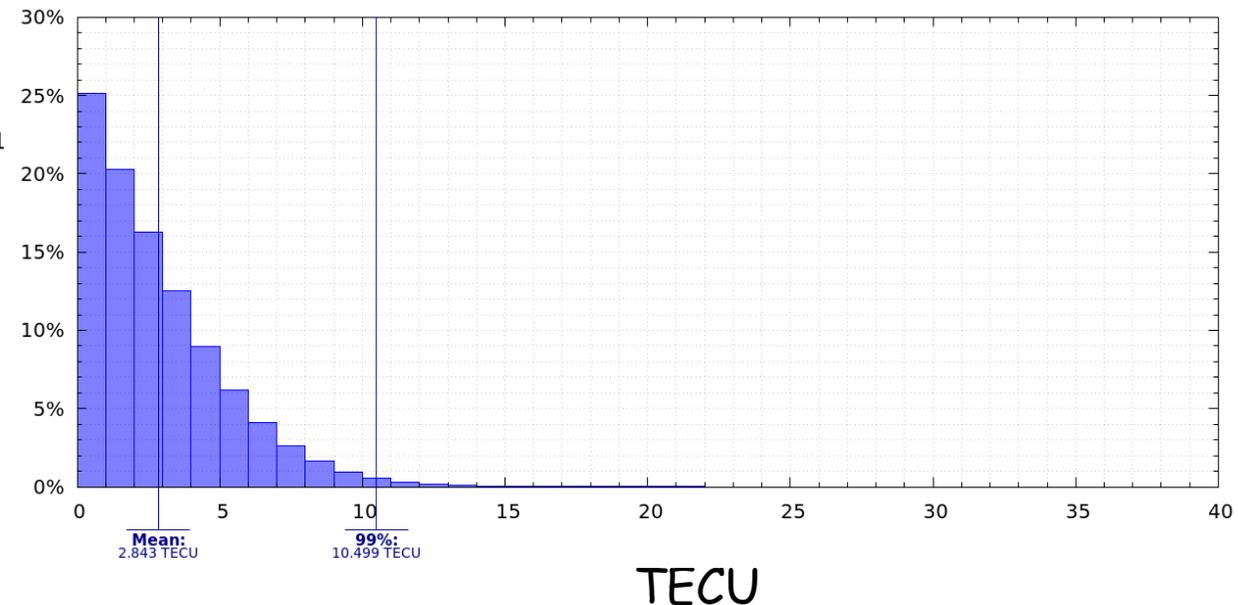


Ionisation gradients - July 2015

Temporal TEC gradient from all the stations in the chosen area

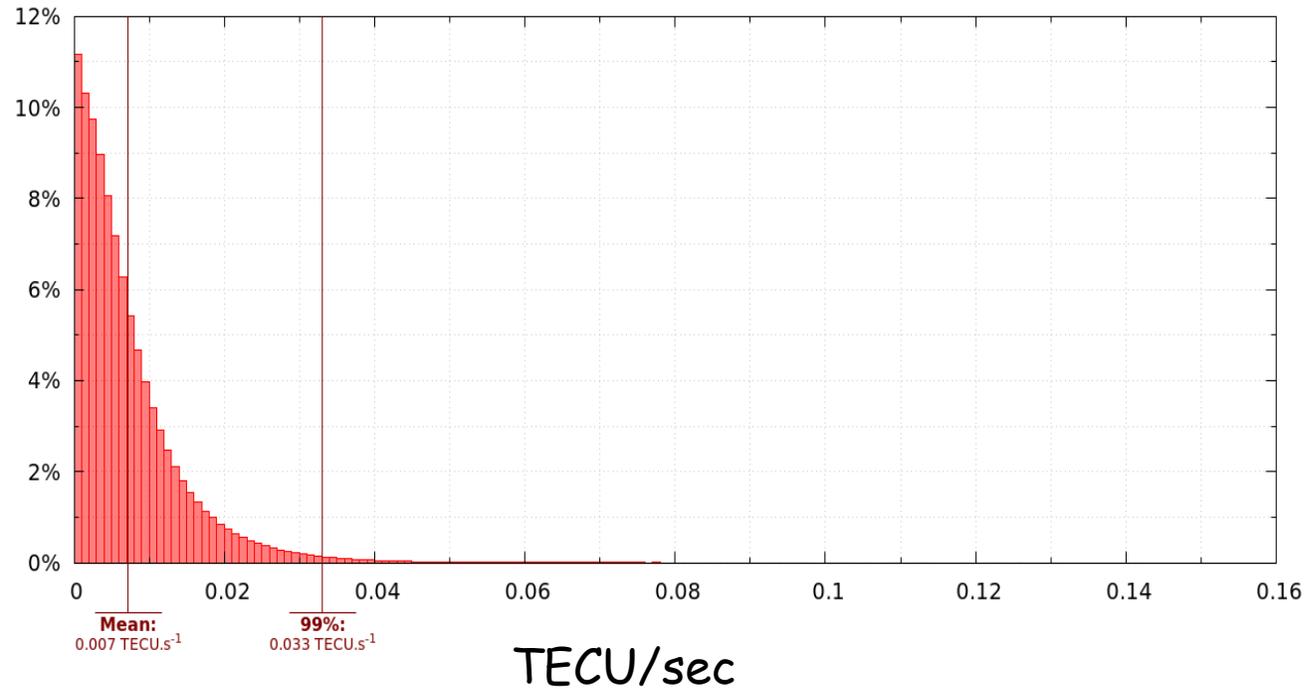


Spatial TEC gradient between two close IPPs

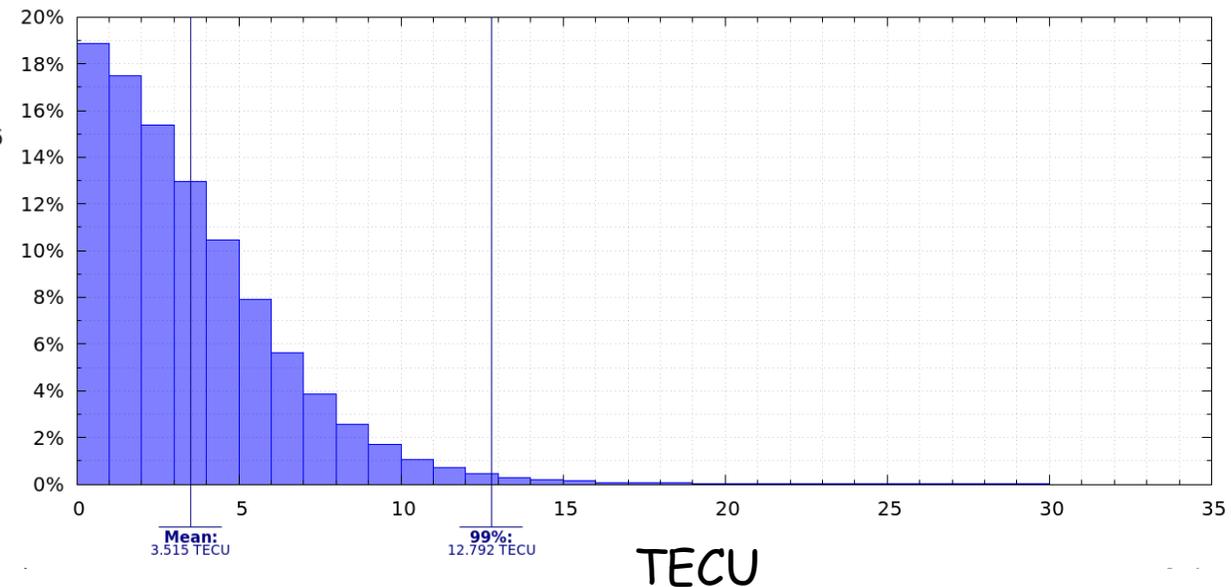


Ionisation gradients - October 2015

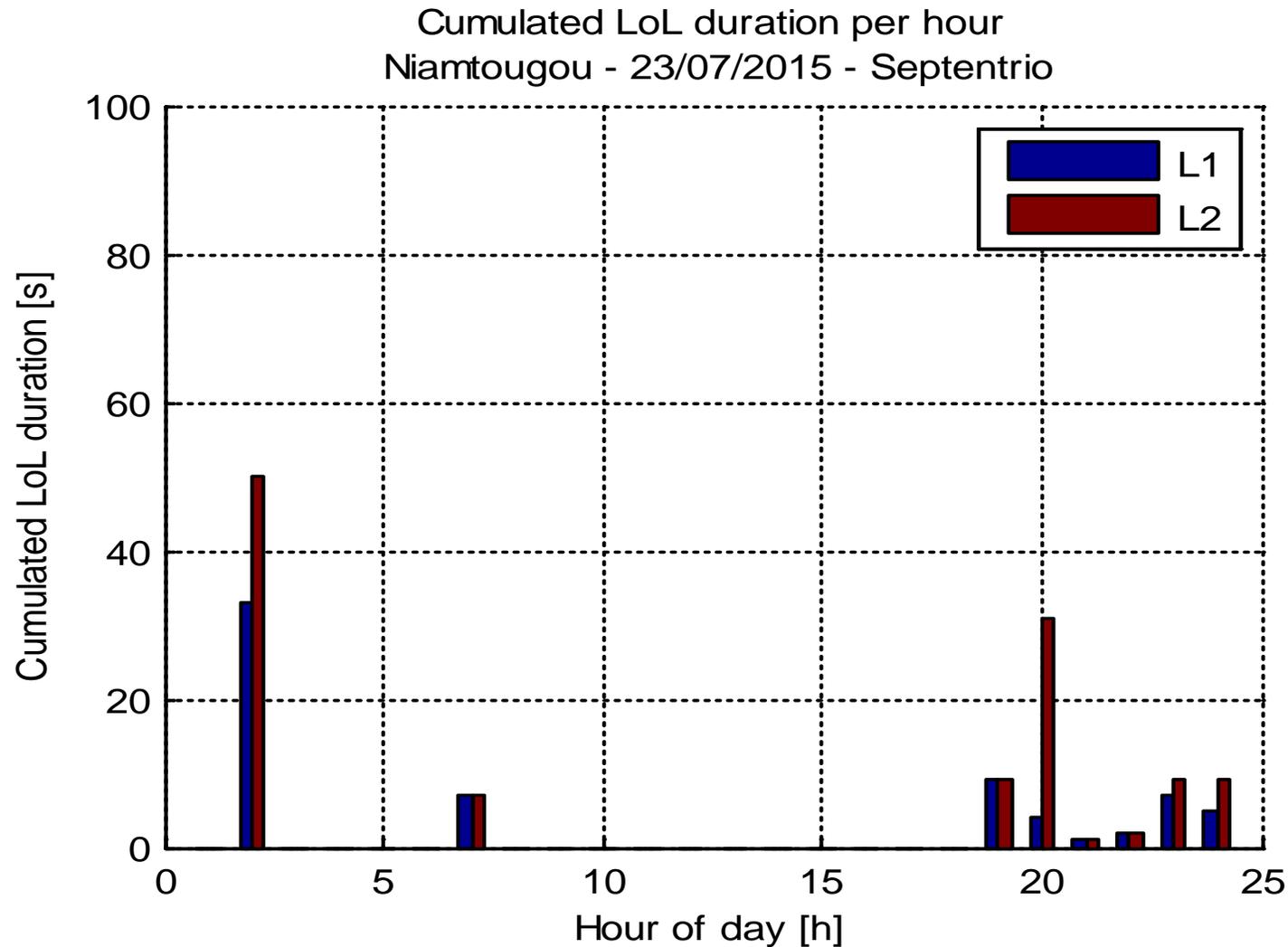
Temporal TEC gradient from all the stations in the chosen area



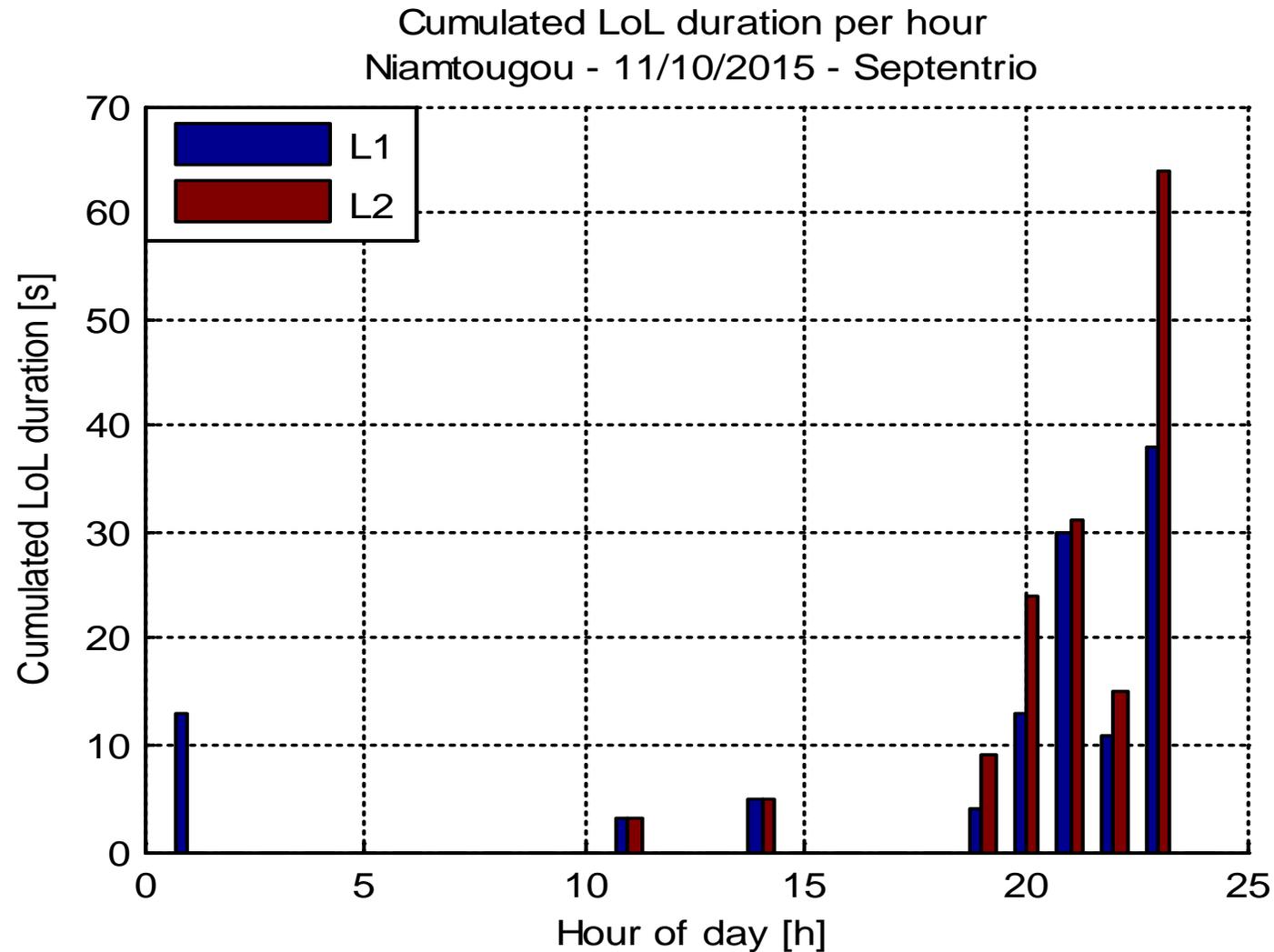
Spatial TEC gradient between two close IPPs



Data gaps introduced by ionospheric scintillation - July 2015



Data gaps introduced by ionospheric scintillation - October 2015



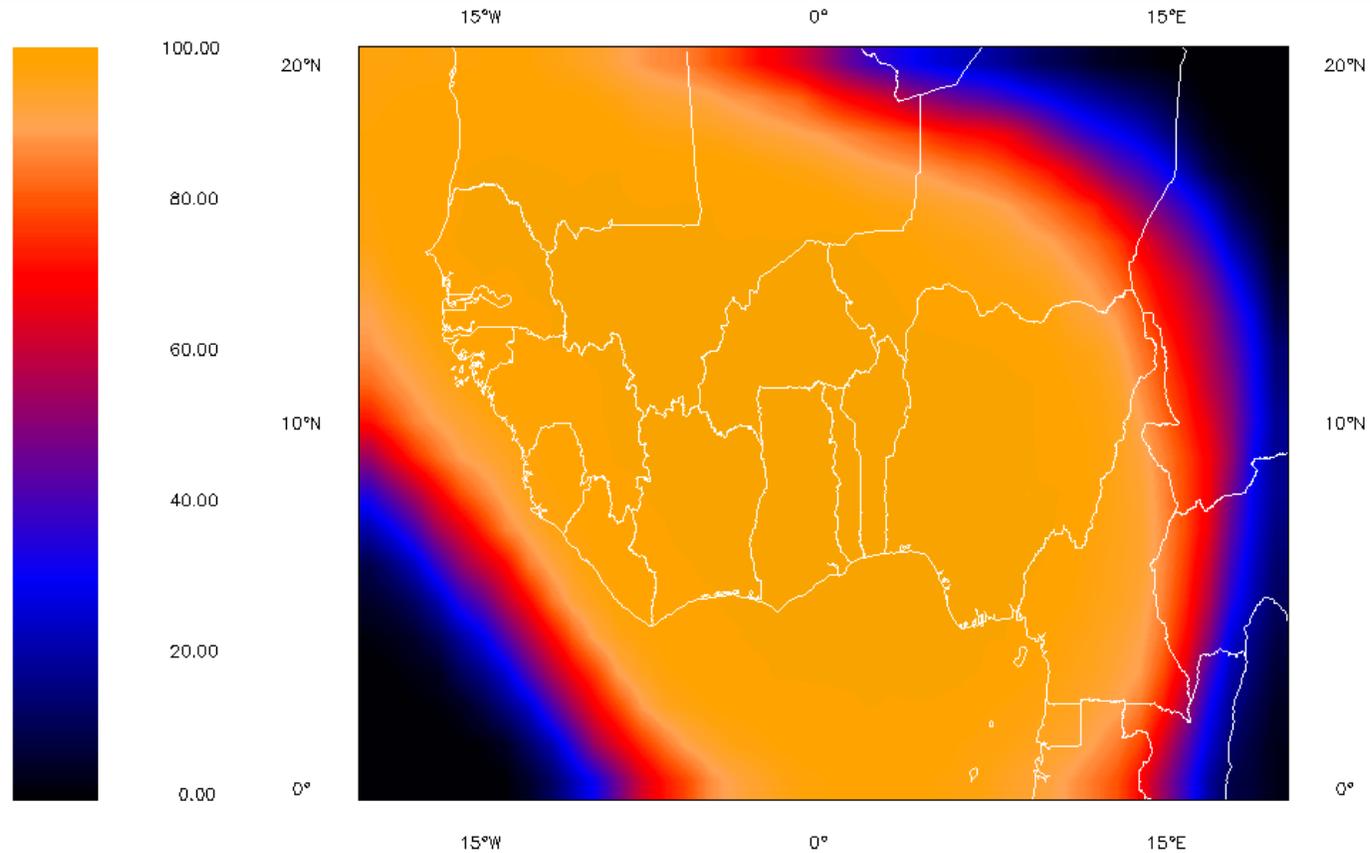
EGNOS availability - July 2015



EGNOS Service Volume Simulator

Average UIM Availability LTLA Map

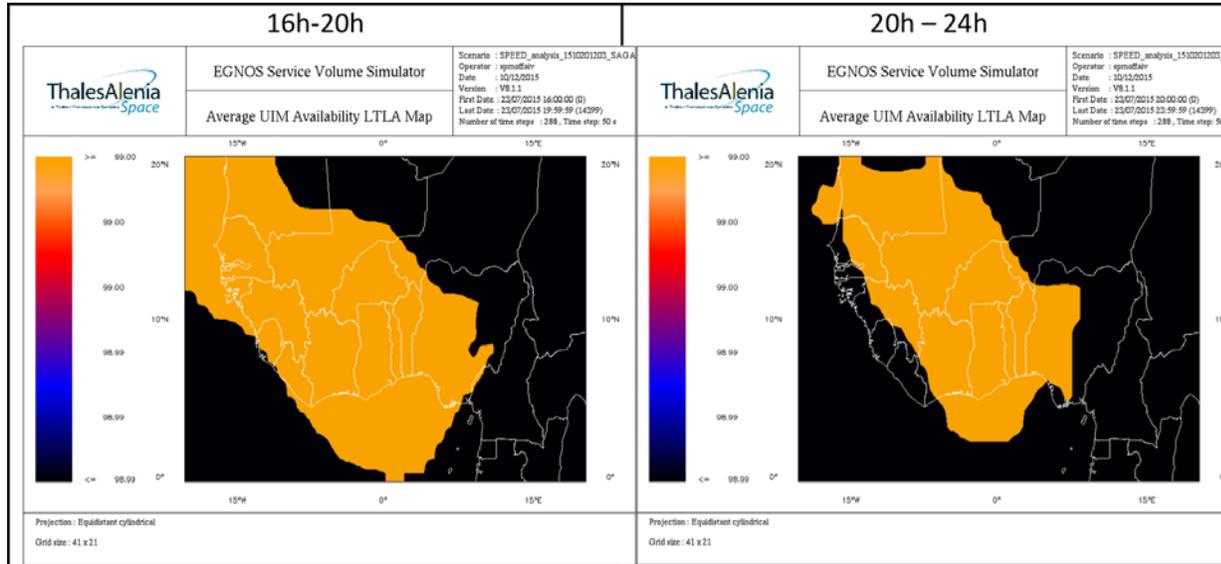
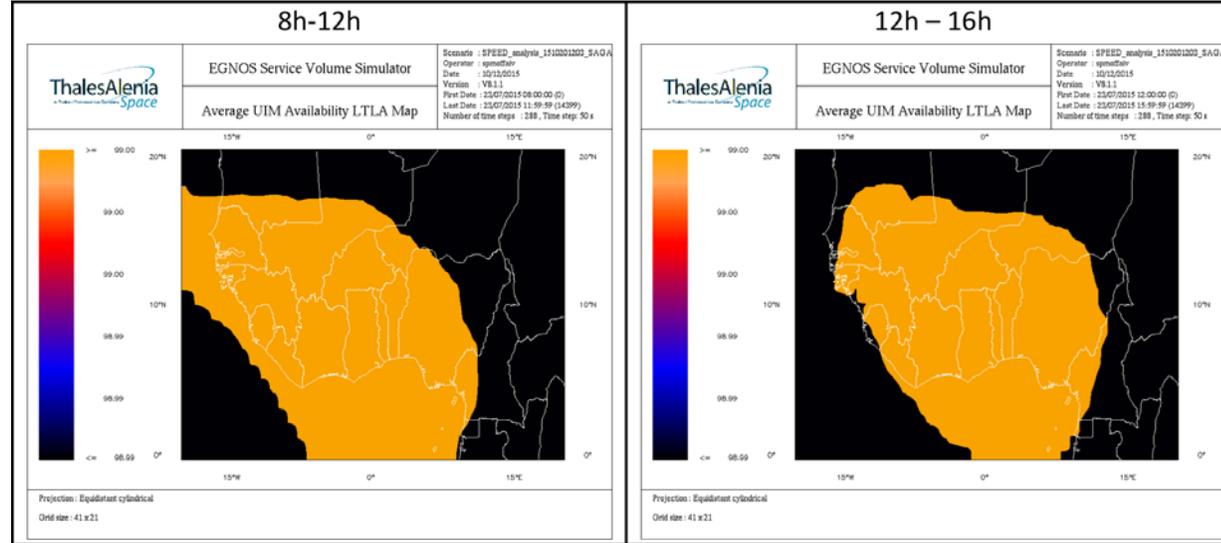
Scenario : SPEED_analysis_1510201203_SAGA
Operator : spmoffaiv
Date : 20/10/2015
Version : V8.1.1
First Date : 23/07/2015 00:00:00 (0)
Last Date : 23/07/2015 23:59:08 (86348)
Number of time steps : 1727 , Time step: 50 s



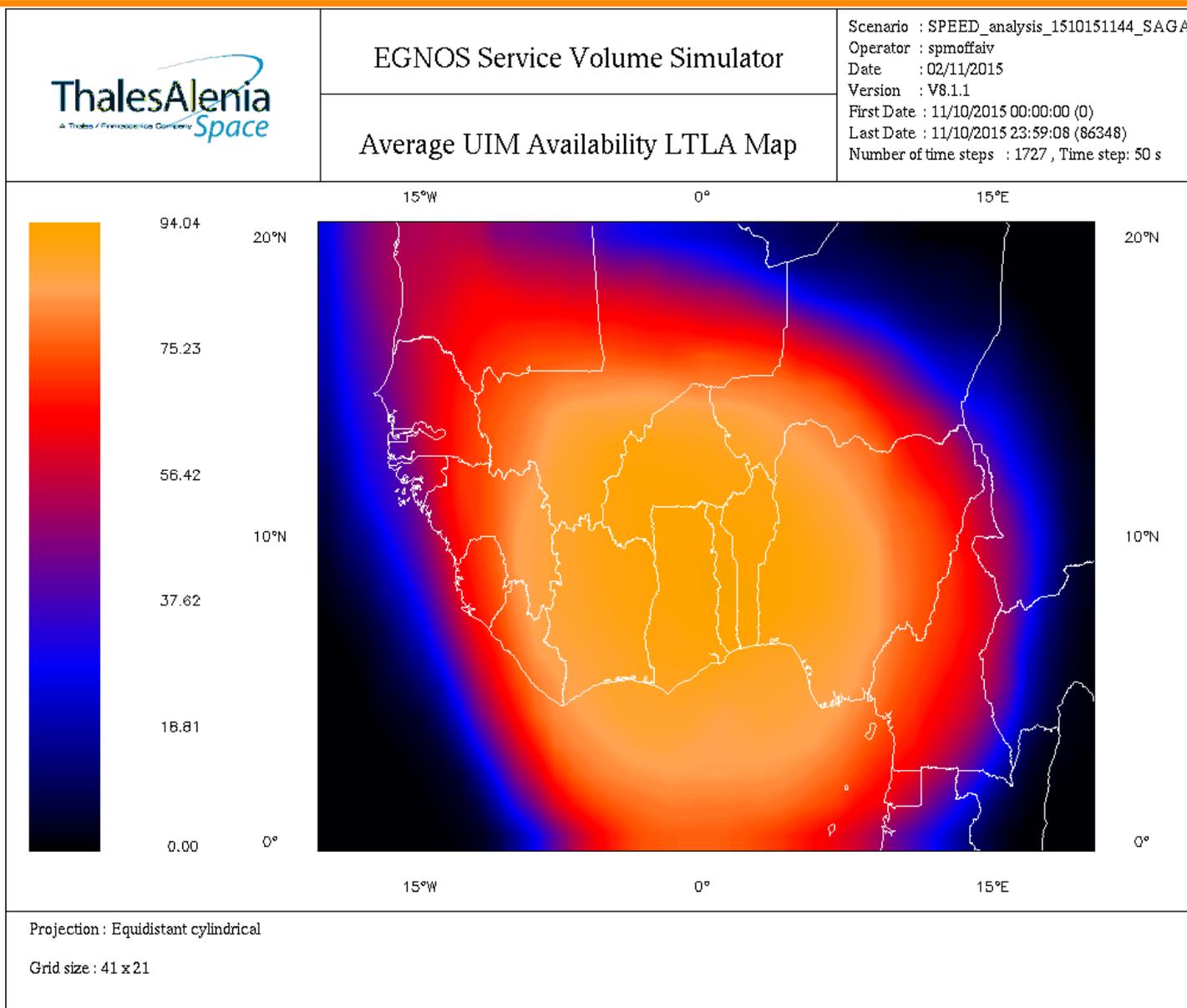
Projection : Equidistant cylindrical

Grid size : 41 x 21

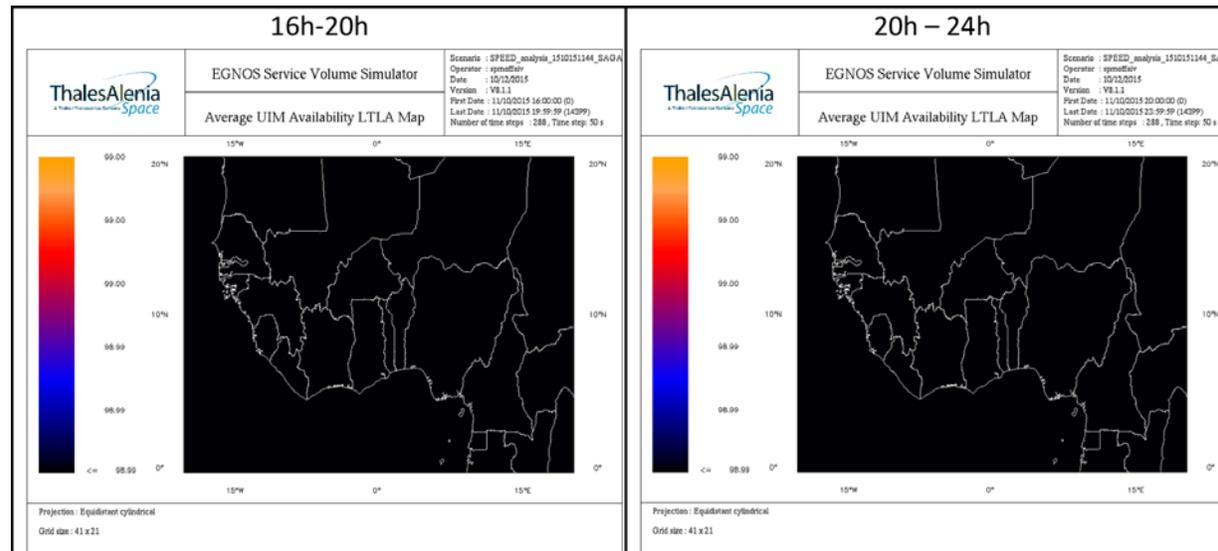
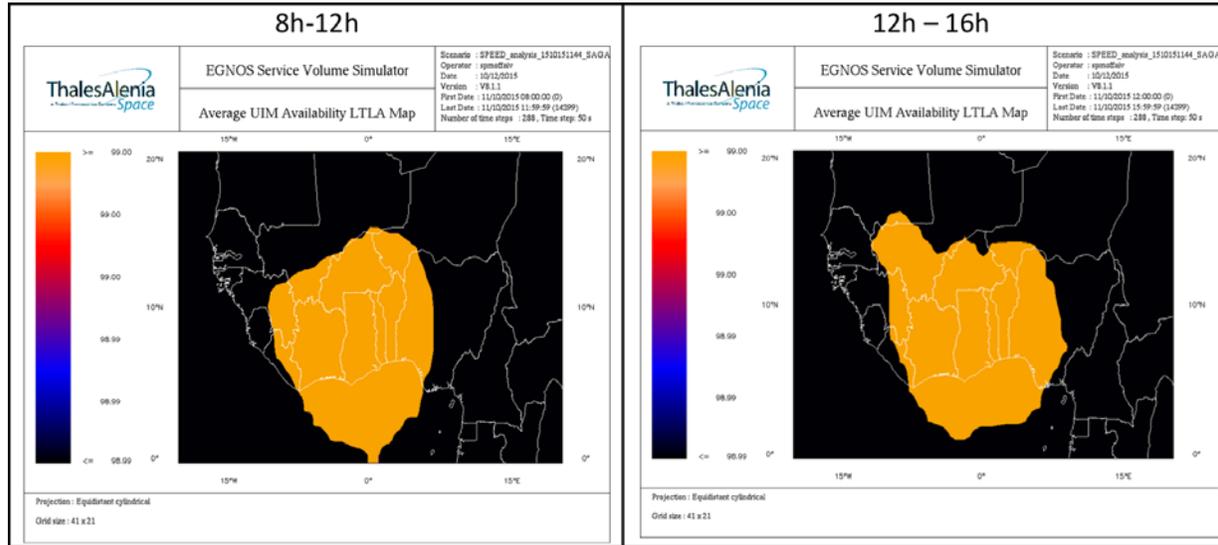
EGNOS availability - July 2015



EGNOS availability - October 2015

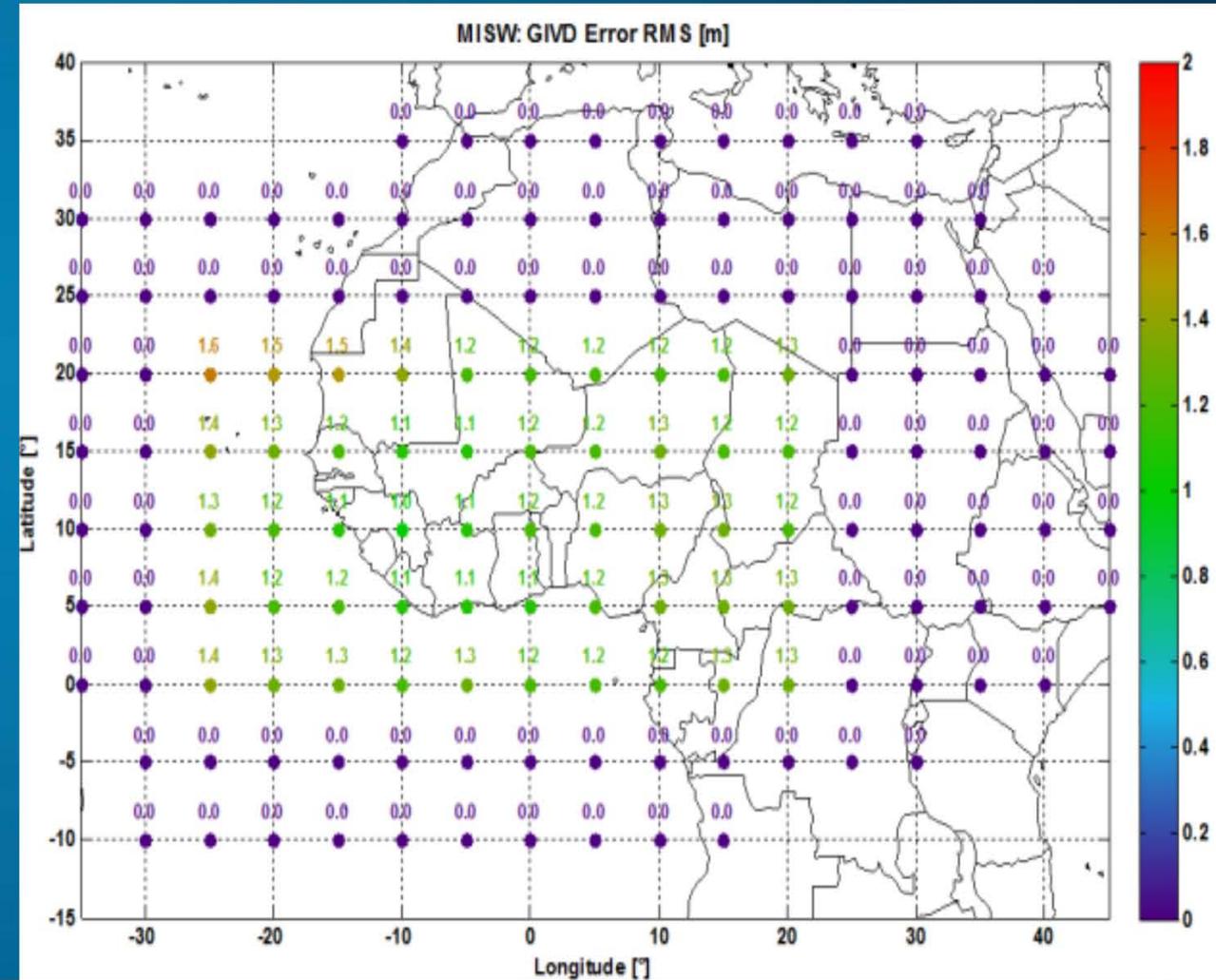
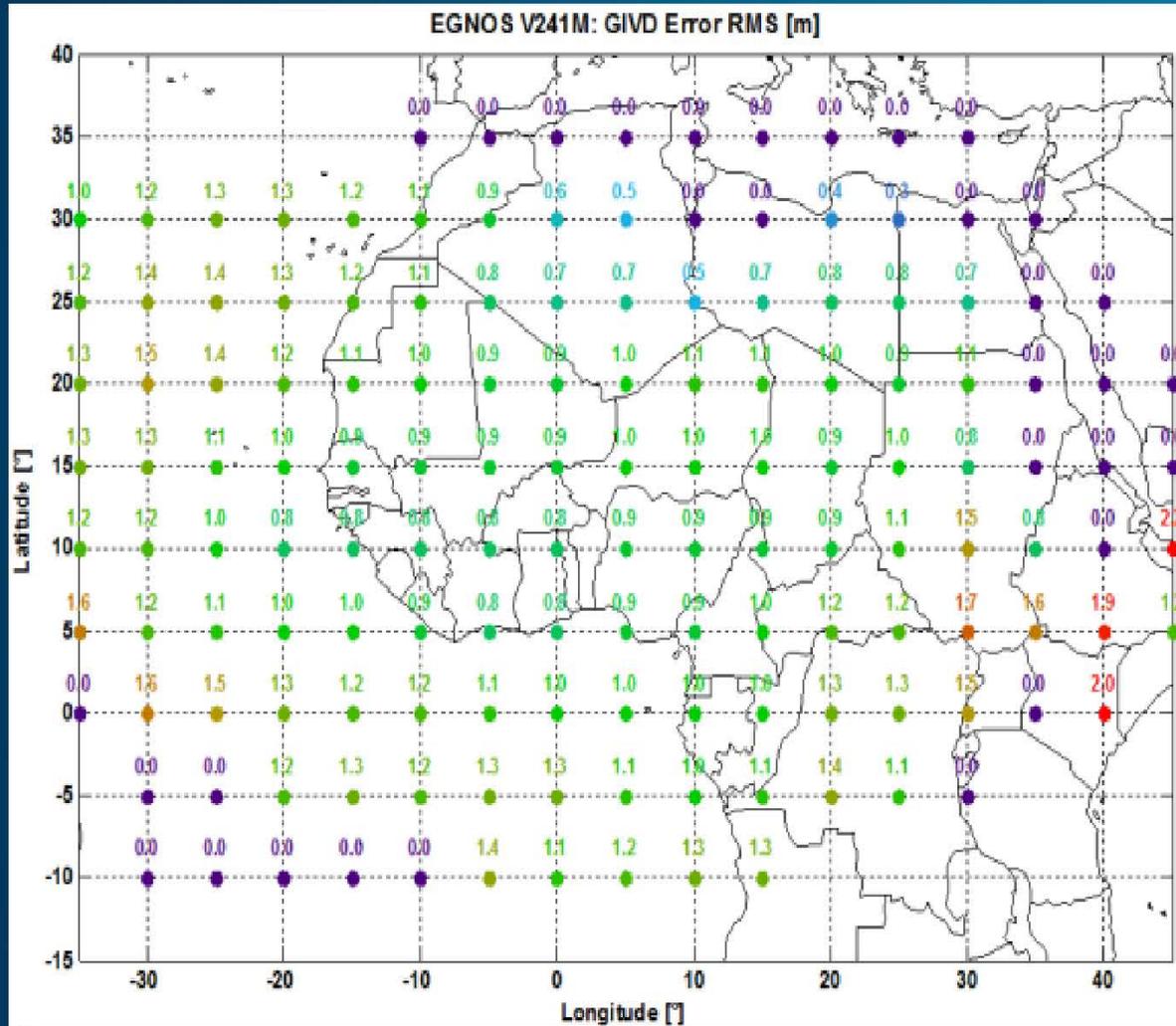


EGNOS availability - October 2015

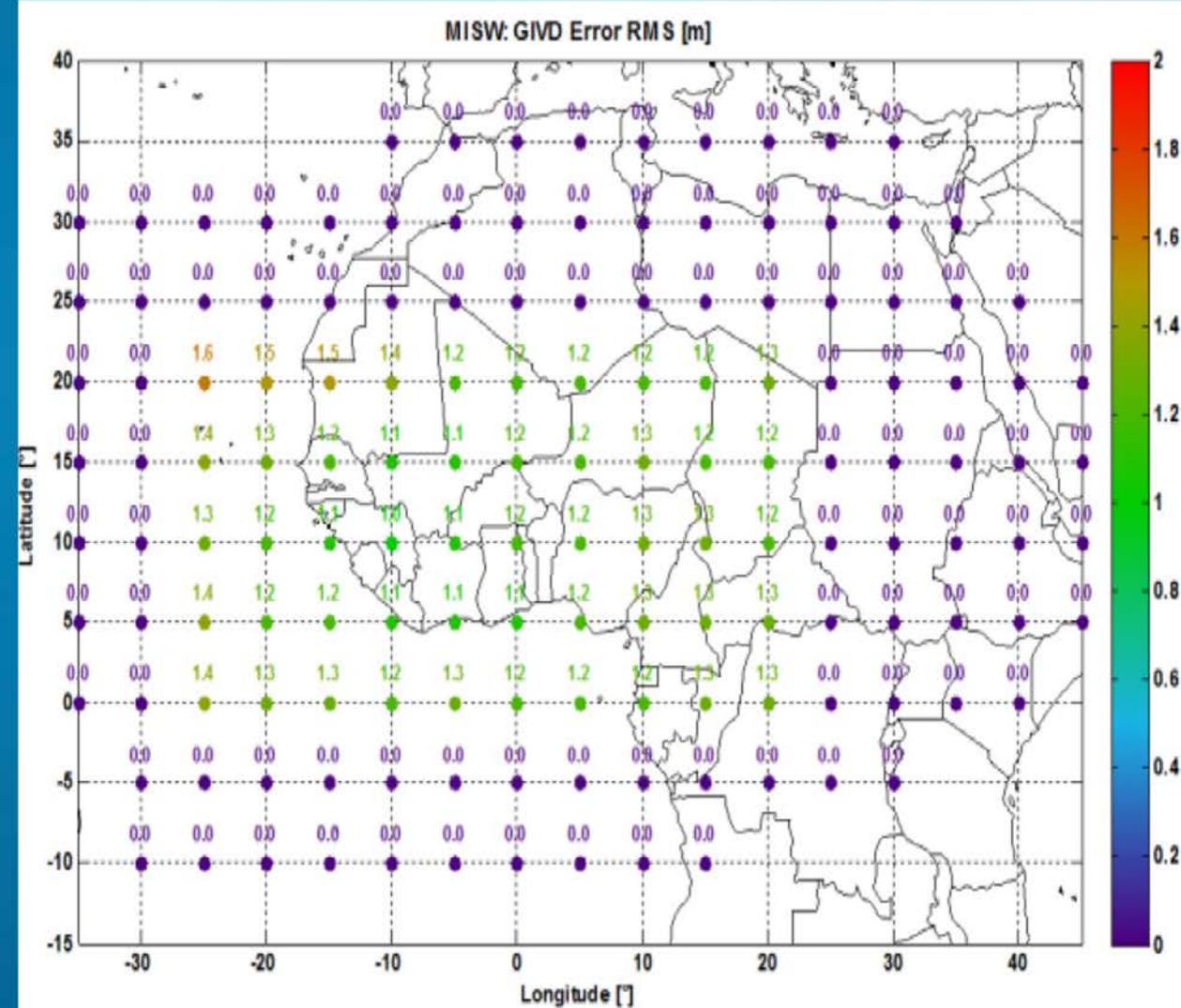
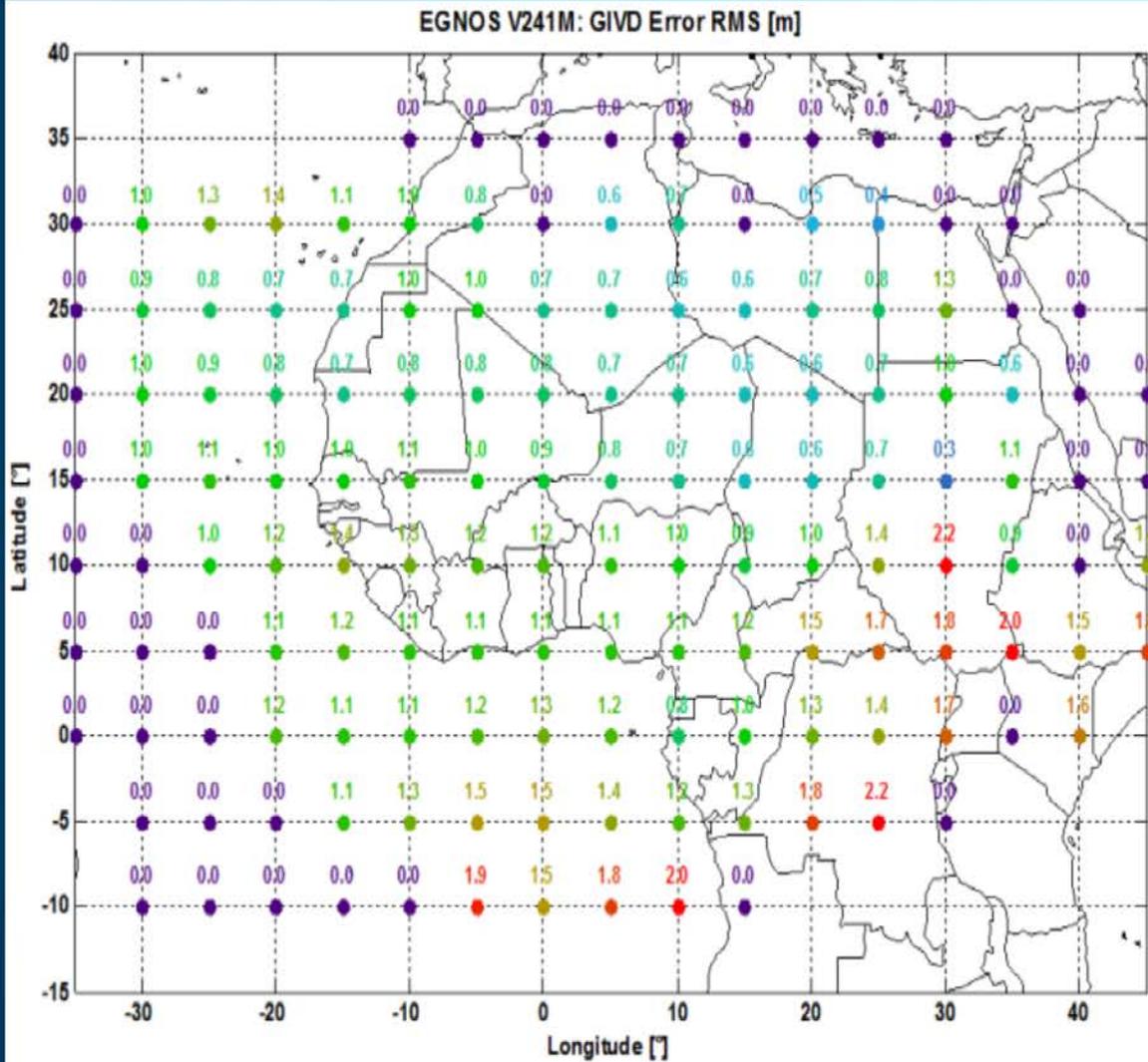


Comparison of different approaches for the grid point correction

Comparison of approaches: 23 July 2015



Comparison of approaches: 11 October 2015



MISW Solutions: Next-generation data-gap-free monitoring station

***GISMO* Prototype**

Robust Carrier Tracking - Simulation Results

Semi-analytic simulator (Matlab)

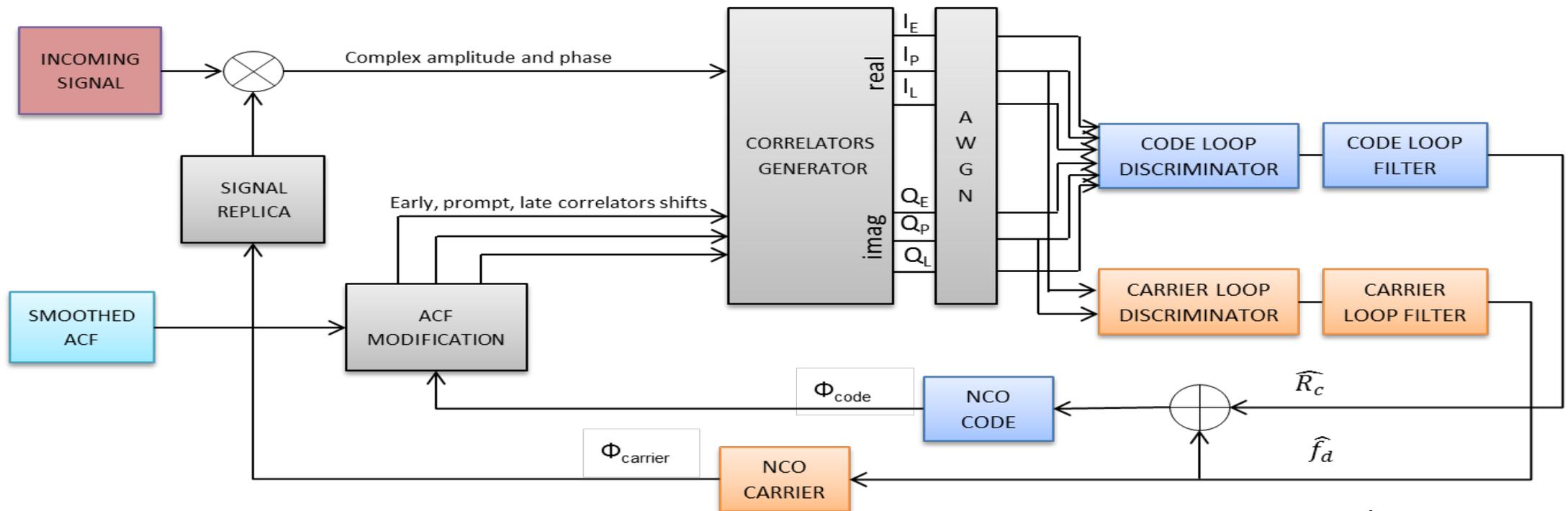
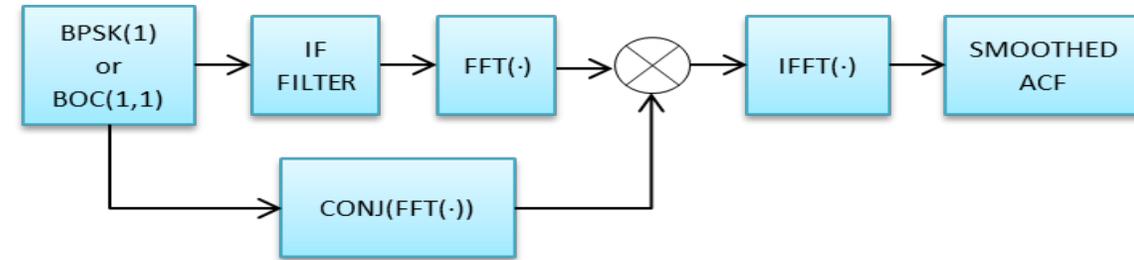
GISM Time series

Techniques comparisons

Semi-analytic Simulator - TASI Receiver Loops Tracking Model

Semi-analytic PLL and DLL combined model

- Simulations at different S4 conditions
→ assess phase and frequency accuracies, robustness vs CS
- Simulation at extreme S4 condition
→ assess robustness vs LoL

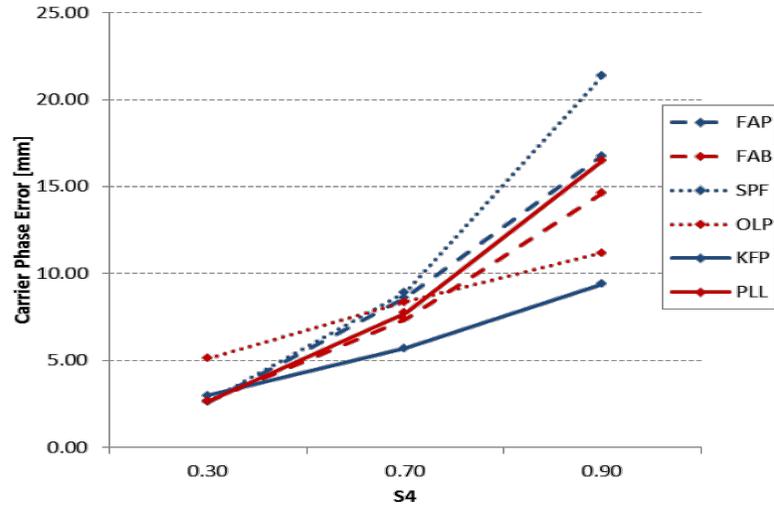


Zin et al, in preparation

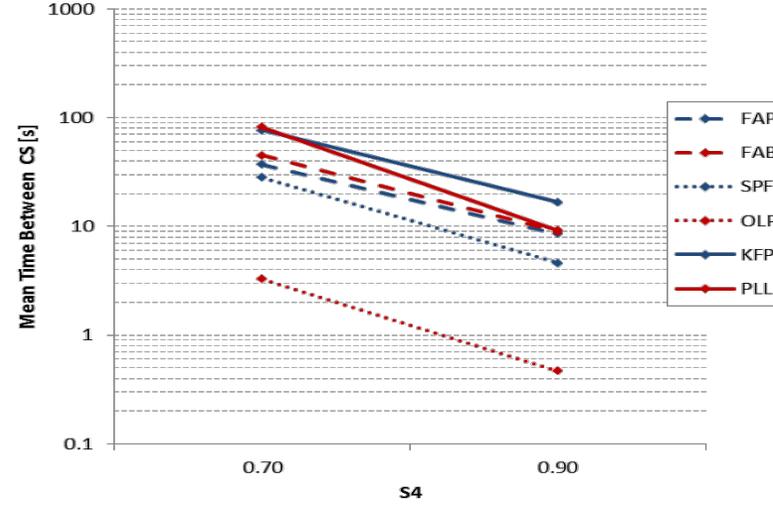
Simulation Results

Zin et al, in preparation

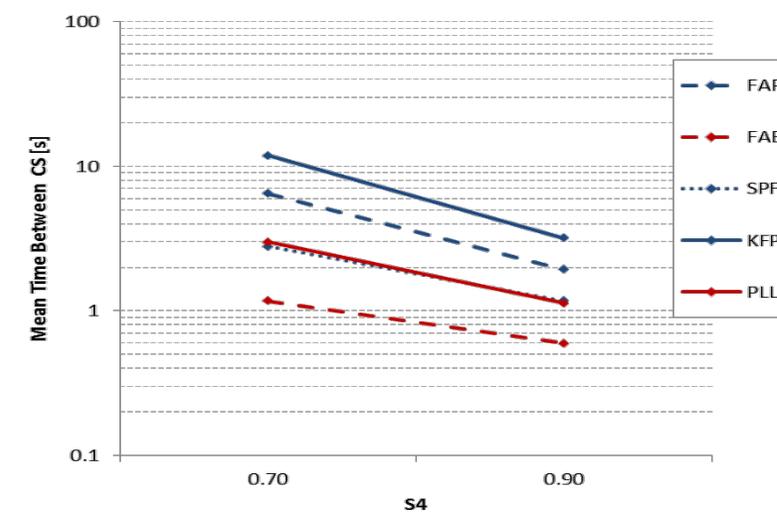
Carrier Phase Error, C/N0 = 30 dB-Hz, atan2 discr.



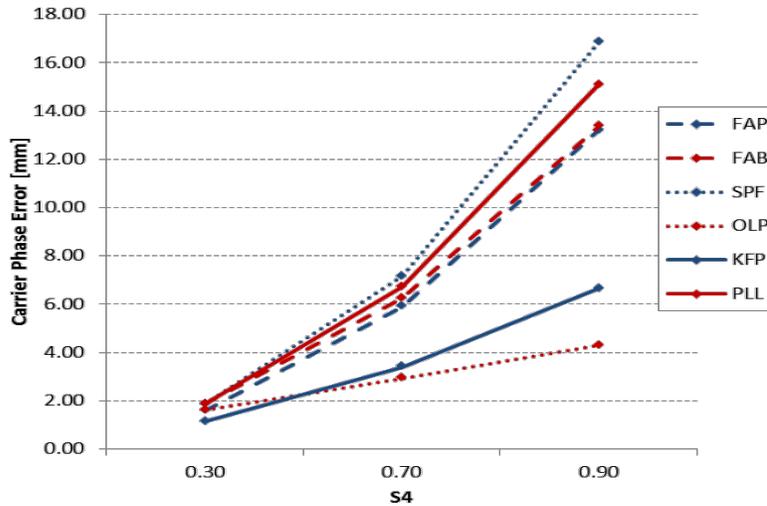
Cycle Slip Occurrences, C/N0 = 30 dB-Hz, atan2 discr.



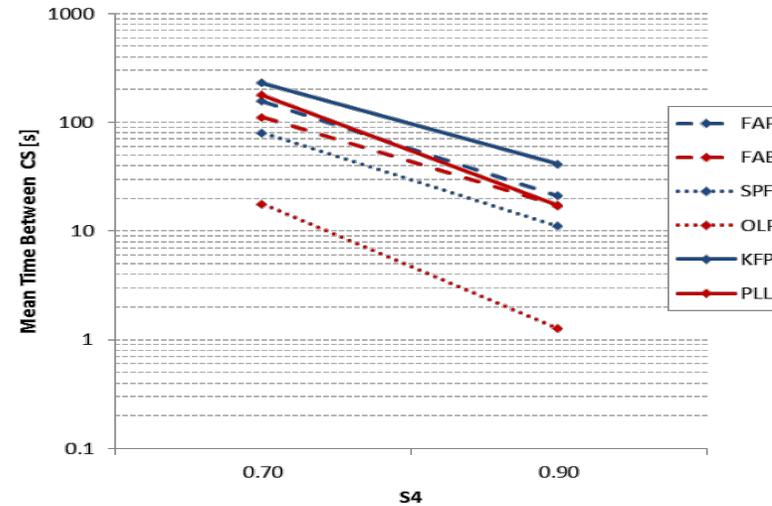
Cycle Slip Occurrences, C/N0 = 30 dB-Hz, atan discr.



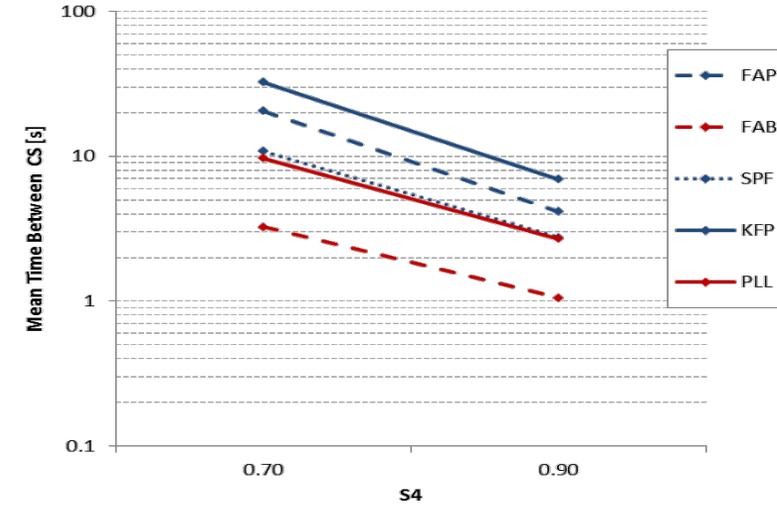
Carrier Phase Error, C/N0 = 40 dB-Hz, atan2 discr.



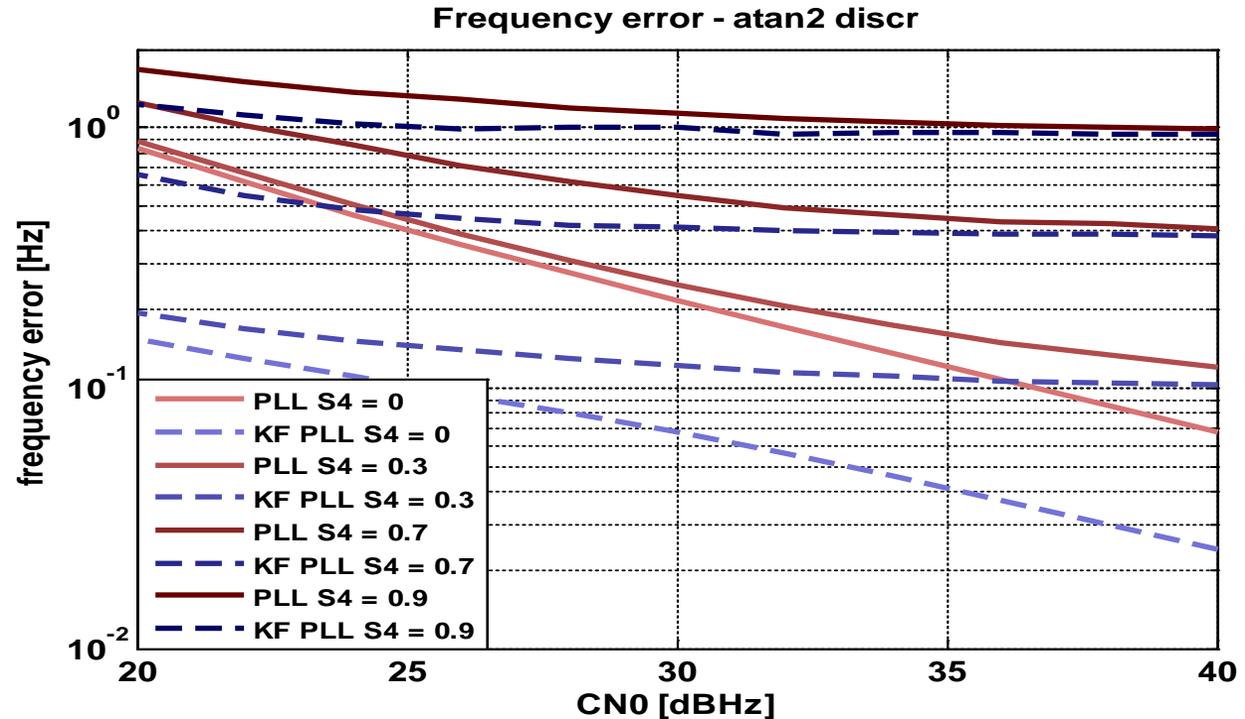
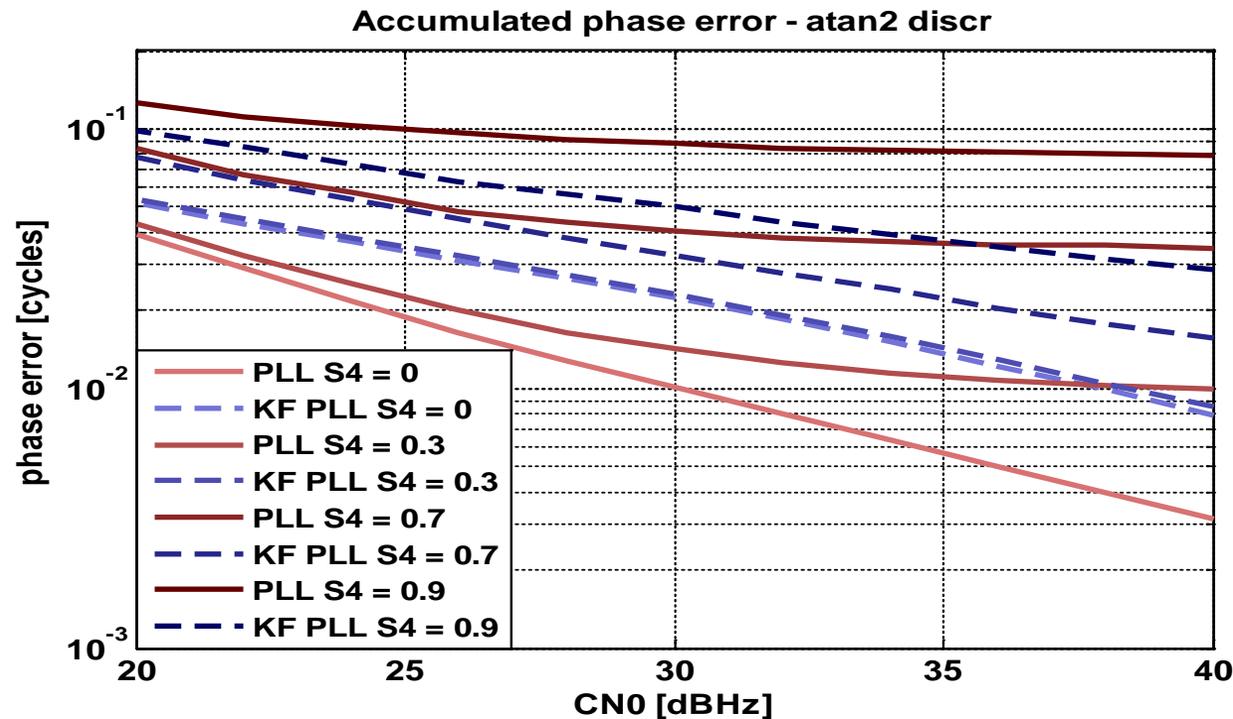
Cycle Slip Occurrences, C/N0 = 40 dB-Hz, atan2 discr.



Cycle Slip Occurrences, C/N0 = 40 dB-Hz, atan discr.



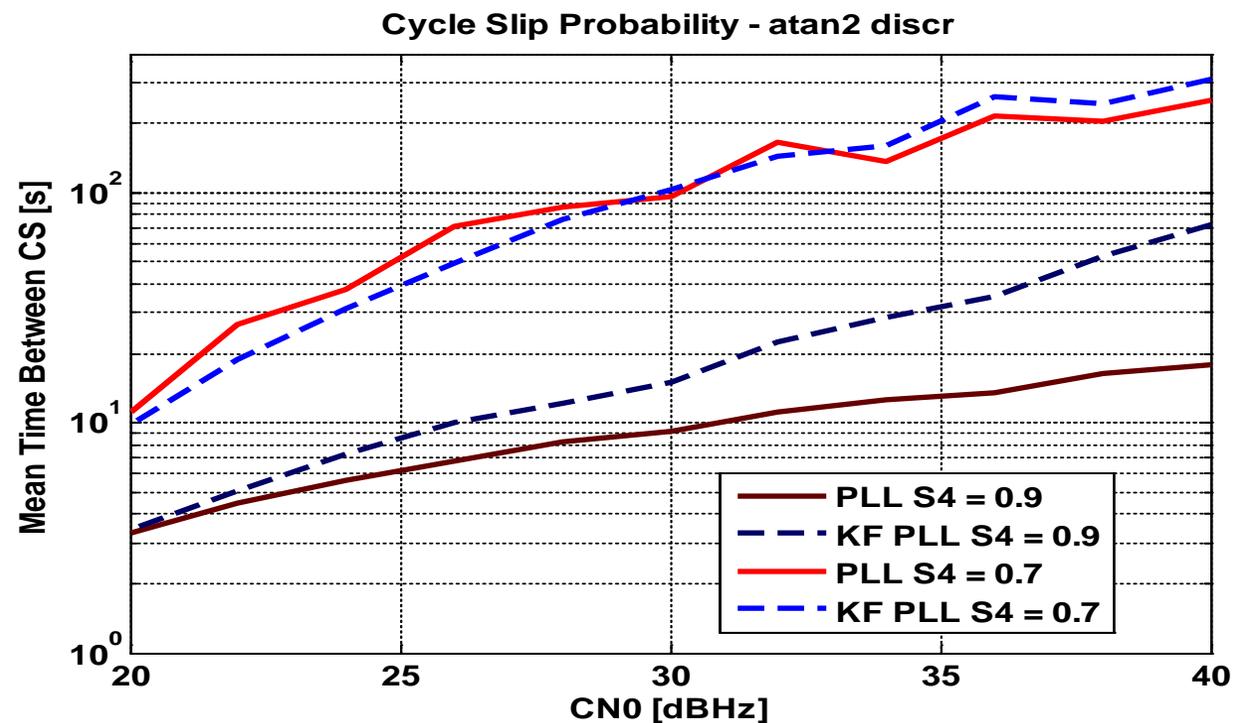
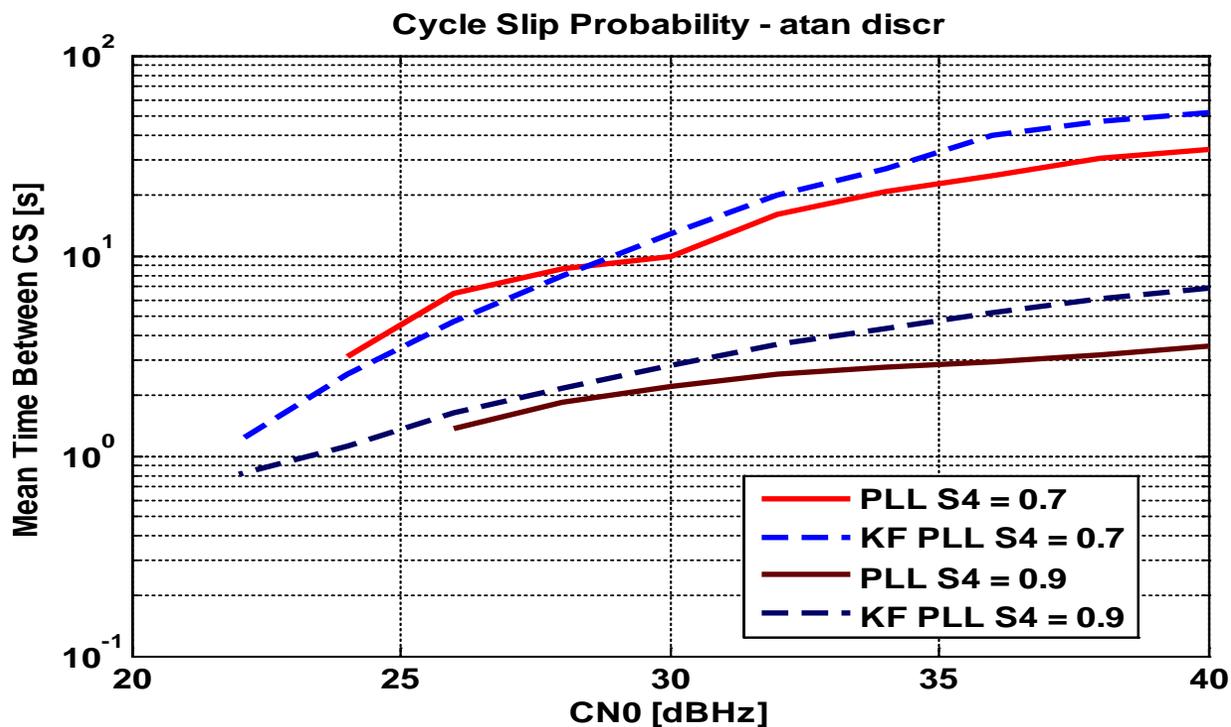
Simulator Results - Phase and Carrier Accuracy



- Kalman filter results to have higher phase errors at lower scintillation, while in presence of strong scintillation it shows a great error reduction, in particular at high C/N_0 values
- The most evident advantage in the Kalman filter reduced frequency error, w.r.t. reference PLL → increasing LoL and frequency False Lock rejection

Zin et al, in preparation

Simultator Results - Robustness vs Cycle Slips



- The figures highlight that for both atan and atan2:
- PLL performs slightly better in case of moderate scintillation,
- a great improvement of KF has been obtained in case of extreme scintillation.

Zin et al, in preparation

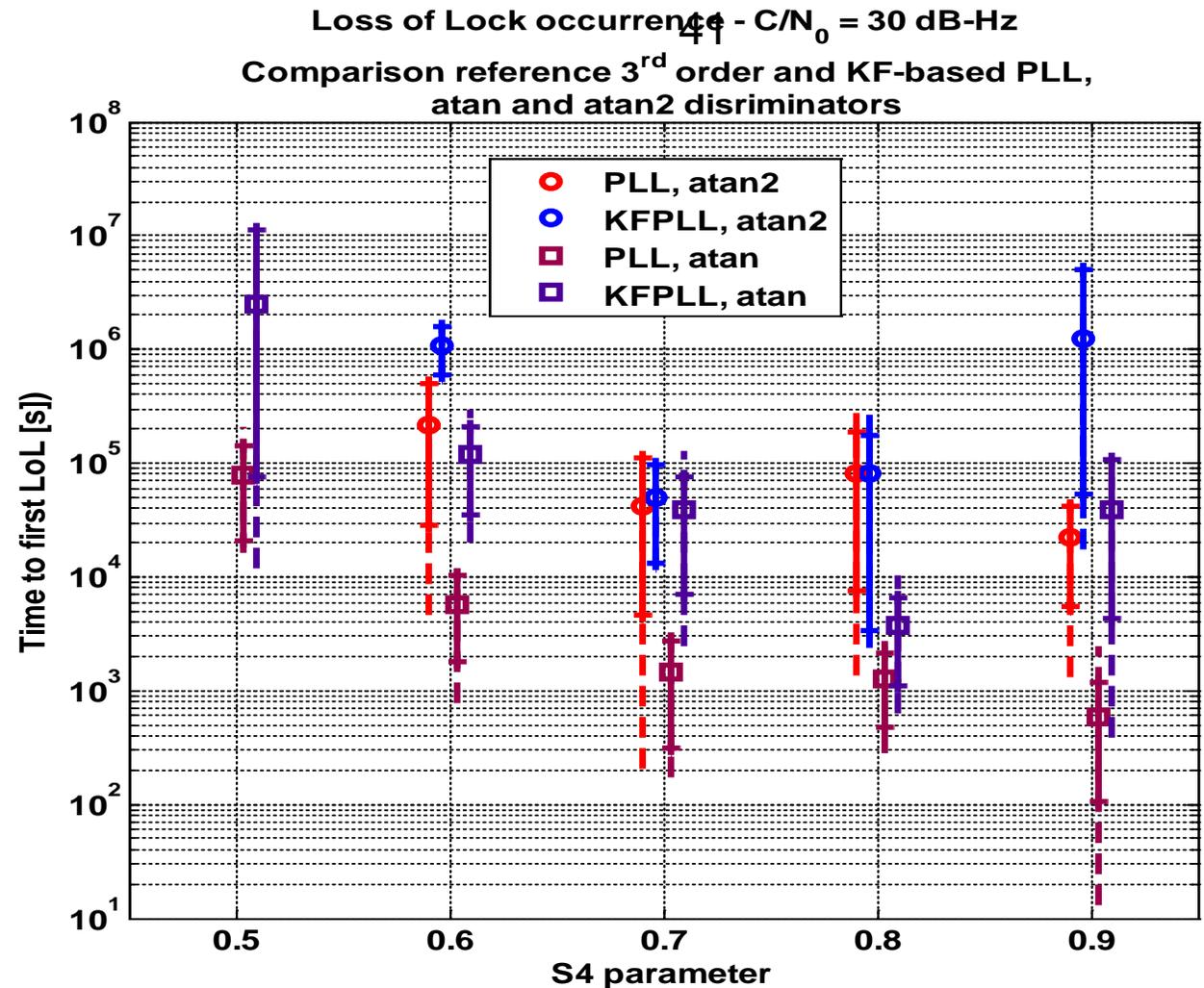
Simultator Results - Robustness vs Loss of Locks

Representation

- Mean: circled points represent average of the collected Times-to-First-LoL.
- Min-Max: vertical lines extend from minimum to maximum Time-to-First-LoL observed.
- The horizontal ticks give representation of the confidence interval

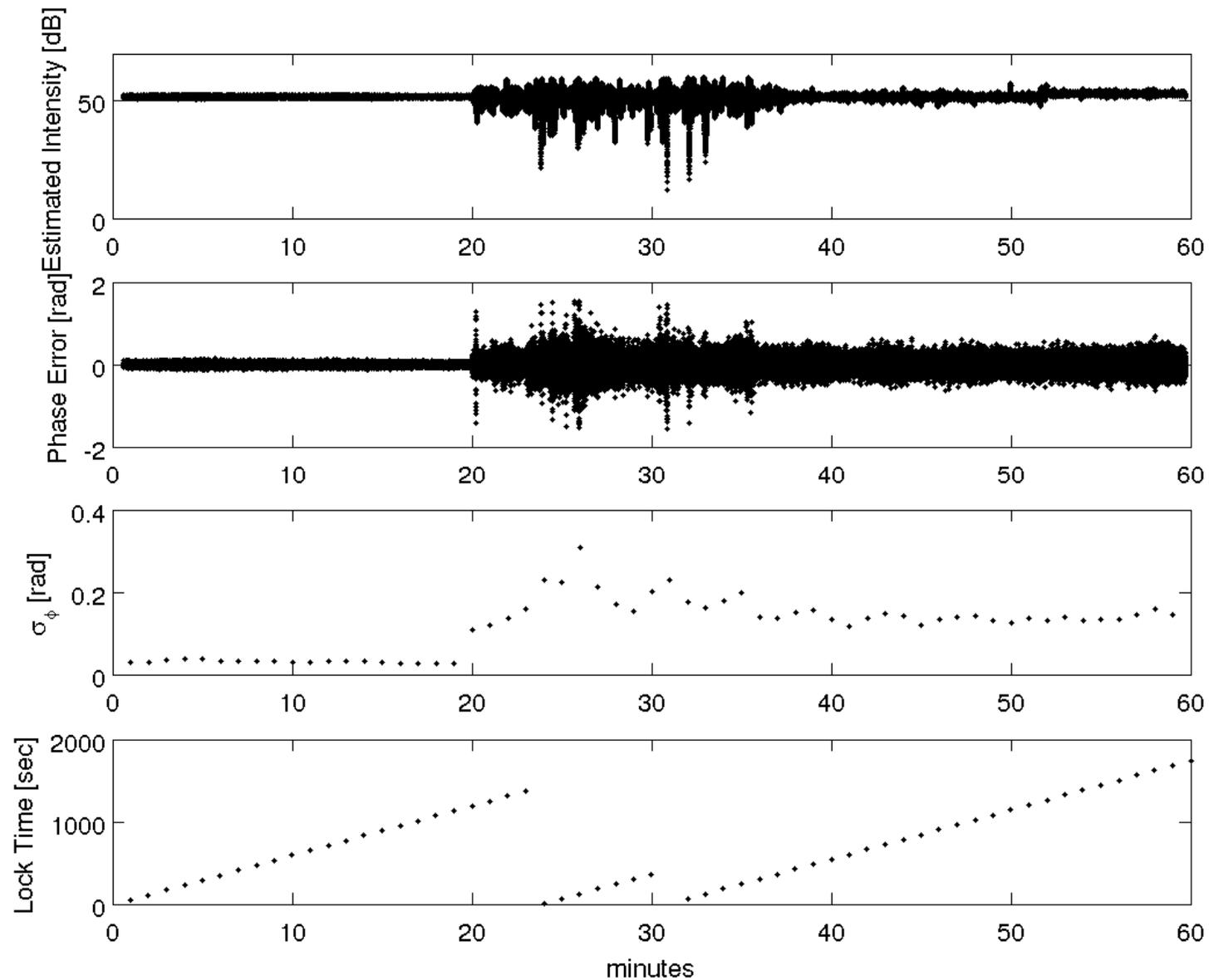
Results

- Benefits brought by KF more evident in atan case
- KF shows increasing robustness in extreme scintillation condition
- atan2 discriminator more robust than atan (no LoL at $S_4 = 0.5$)



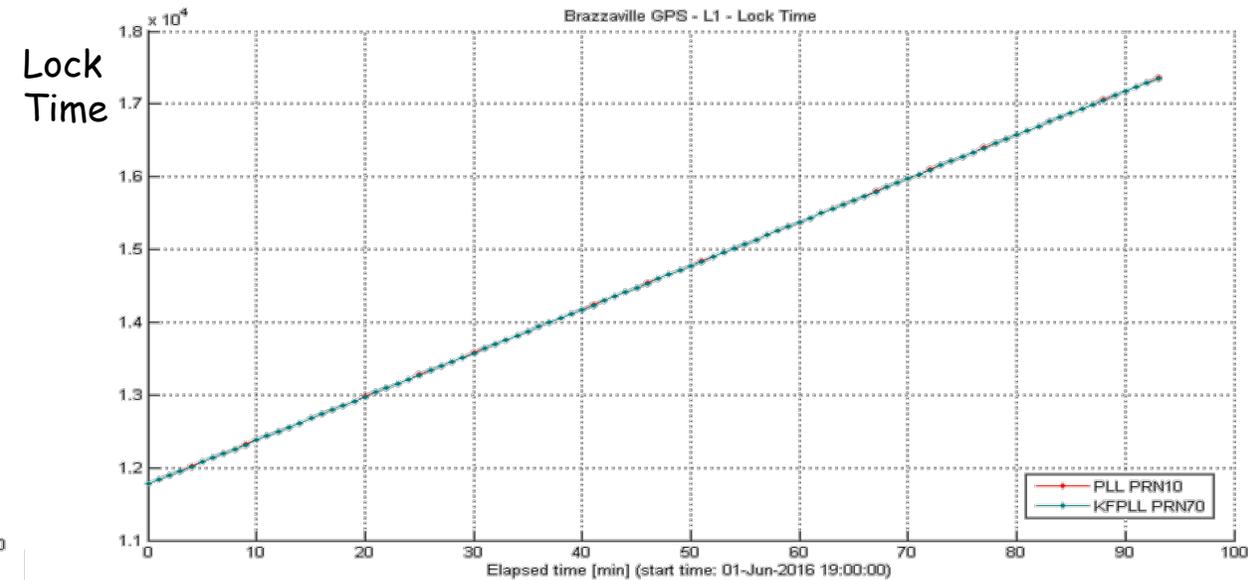
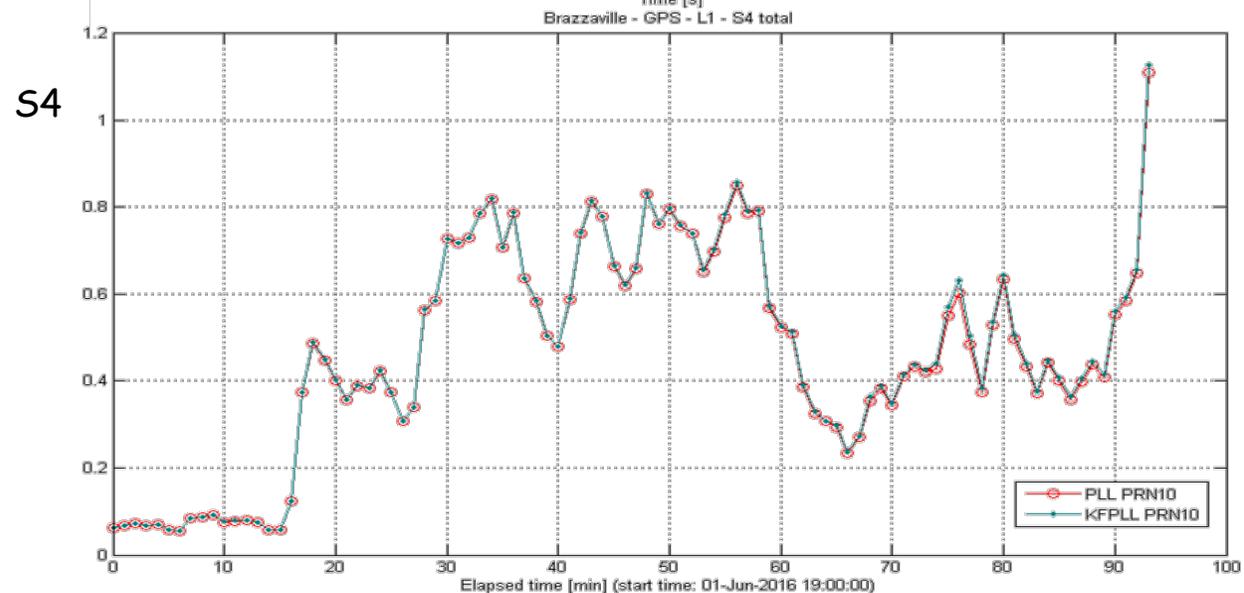
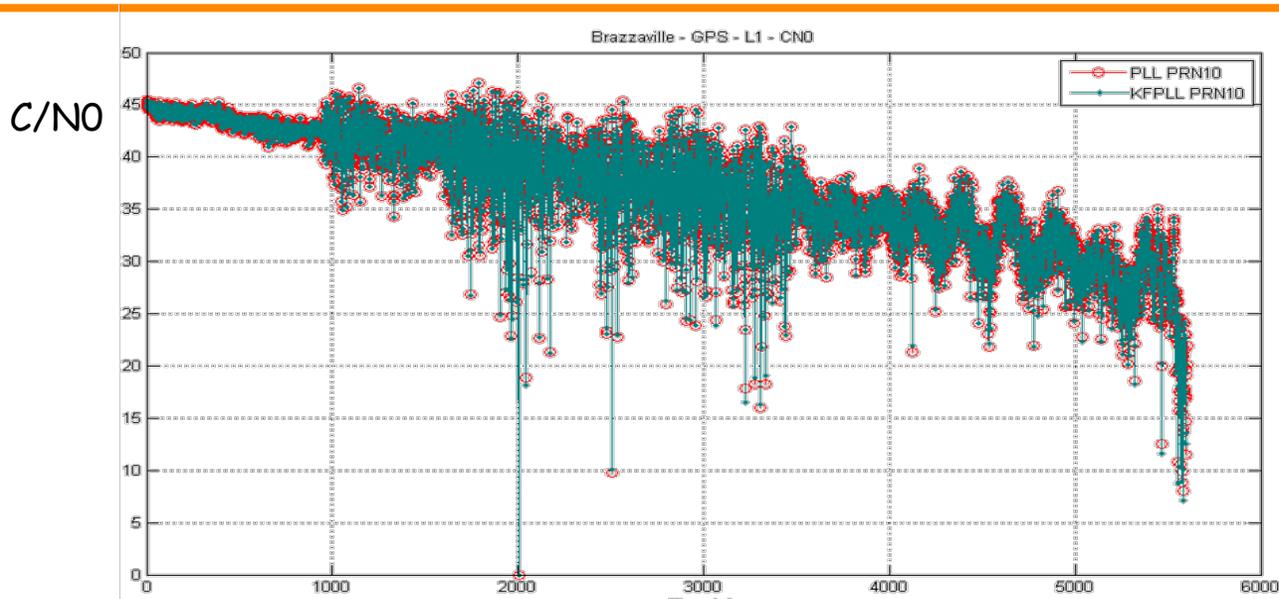
Zin et al, in preparation

Additional tests (on-going): perturbations from real data



(Forte, 2012)

Updated GISMO Receiver Live Campaign



- Data recorded in May/June 2016 by updated GISMO RX show ionospheric scintillation activities
- Same scintillation level recorded in September/October 2015 which led to LoL
- No LoL observed for both carrier tracking strategies with new SW version

Conclusions

- Multi-constellation multi-frequency monitoring station
- All Kalman filter based PLL channels
- Intelligent Loss of Lock indicator
- Improved L2C tracking through L1CA-frequency aiding
- Improved acquisition time on GPS L5, Galileo E1BC and E5a
- Robustness in the presence of both low and high latitudes conditions
- On-going development of system algorithms



Thank you for the attention

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